

Theoretical Rotational Spectra of Deuterated Benzene Isotopologues: Improvements on the Benzene's quantification at Interstellar Medium

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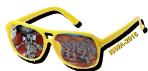
*Centro Federal de Educação Tecnológica de Minas Gerais

†Universidade do Vale do Paraíba

‡Universidade Federal de Minas Gerais

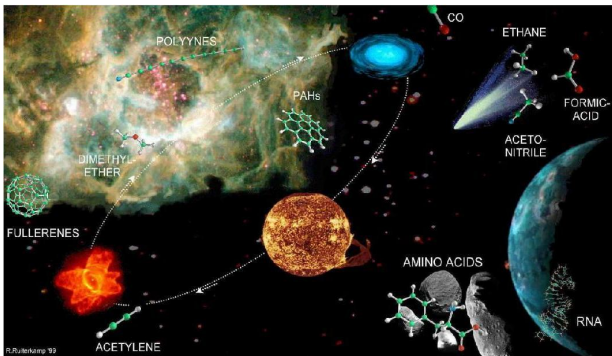
July 8, 2016

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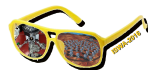


Molecules at ISM and CSS

About **195** molecules detecteds in the Interstellar (**ISM**) medium and Circumstellar Shells (**CSS**) (as of 06/2016)².



²The Cologne Database for Molecular Spectroscopy - CDMS, H. S. P. Muller et. al., J. Mol. Struct. 742, 215-227 (2005)

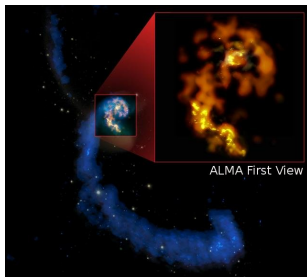


Molecules at ISM and CSS

Latest generation instrumentation to **Radioastronomy**.

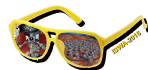


(a) ALMA Radiotelescope



(b) NGC 4038 and 4039

ALMA - *Atacama Large Mm/Submm Array*
64 antennas at Chajnantor plateau - Chile (5058,7 m)
Front End: 10 bands between 31 - 950 GHz



Motivation: Deuterated isotopologues at ISM

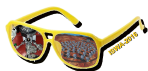
- Deuterated Polycyclic Aromatic Hydrocarbons **PAD** in **ISM**².
- Key role of **PAH** in the astrochemical evolution of the **ISM**
- Formation of benzene in the **ISM**³.

Proposal of Detection

The search for **deuterated isotopologues** of otherwise apolar molecules can be amplified using radio-telescopes if we have good **synthetic spectra** for these ones.

²Peeters et al., *Astrophys. J.*, 604, 252 (2004)

³Jones et al., *Proc. Nat. Ac. Soc.*, 108, 452 (2010)



Quantum Chemical Calculations

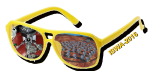
- Theoretical and computational approach for building highly accurate *ab initio* **pure rotational spectra** of deuterated isotopologues.
- Using these spectra to assist possible detections of polyatomic isotopologues of astrophysical interest.⁴⁵.

Simulating **Pure Rotational Spectra**

- Level of approximation: our pure rotational spectra simulations needs **dipole moments**, **rotational constants**.
- We will not to consider **centrifugal distortion constants**, **Hyperfine** and/or **Spin** effects.

⁴R. Motiyenko, J. Phys. Chem. A, 2015, 119 (6),1048-1054

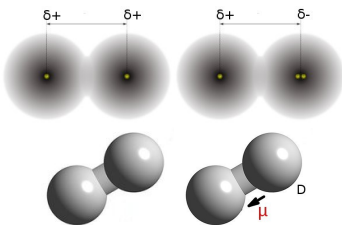
⁵C. Puzzarini, Phys. Chem. Chem. Phys., 2013,15, 6595-6607



The Isotopic Effect on Dipole Moments

Molecular Symmetry Breaking

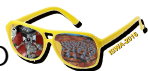
e.g.: Hydrogen atoms (H) replaced by deuterium atoms (D)



Isotopic asymmetry on the **electronic** distribution

$$a_0 = \frac{4\pi\epsilon\hbar^2}{me^2} \quad (1)$$

Consequence: dipole moment (DM) $\sim 10^{-3}$ debye to HD



Electronic Structure - Molecular Approximations

$$\hat{H}\psi(\vec{r}, \vec{R}) = E\psi(\vec{r}, \vec{R}) \quad (2)$$

Adiabatic Model: Born-Oppenheimer Approximation(BOA)

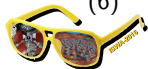
$$\hat{H} = \hat{T}_{nuc} + \hat{H}_{BO} \quad (3)$$

$$\hat{H}_{BO} = - \sum_i \frac{\nabla_i^2}{2} + V \quad (4)$$

$$\hat{H}_{BO}\phi(\vec{r}; \vec{R}) = E_{ele}\phi(\vec{r}; \vec{R}) \quad (5)$$

$$E_{tot} = E_{ele} + \sum_{A=1}^M \sum_{B>A}^M \frac{Z_A Z_B}{R_{AB}} \quad (6)$$

Potential Energy Surface - PES



Electronic Structure - Molecular Approximations

The **BOA** can't to predict the isotopic effect



FNMC - Finite Nuclear Mass Correction

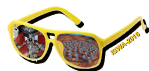


The FNMC Model Hamiltonian FNMC^a

^aJ.R. Mohallem, J. Phys. B: At. Mol. Opt. Phys. 32 (1999) 3805

$$\hat{H}_{FNMC} = \sum_A^m \left(- \sum_i^n P_A \frac{\nabla_i^2}{2M_A} P_A \right) + \hat{H}_{BO} \quad (7)$$

The computational cost is identical to the **BOA**



Ab Initio Computational Packages Adapted to FNMC

- **GAMESS/ISOTOPE** - Gonçalves and Mohallem - 2003⁶
- **deMon 2k** - Mohallem et al - 2008⁷

DaltonFNMC - 2011^a

^aArapiraca et al., J. Chem. Phys. 135, 244313 (2011)

- Dr. **Dan Jonsson** (CTCC, University of Tromsø- Norway)
- Dalton 2.0/2016: *Ab Initio* and *DFT* methods
- www.daltonprogram.org - Open Source Code

⁶C. P. Goncalves and J. R. Mohallem., J. Comp. Chem., 25, 1736, (2004)

⁷J. R. Mohallem et al, J. Phys. Chem. A, 118, 8896, (2008)



Vibrational Corrections on Molecular Properties

Theory X Experiment

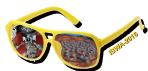
Computing **nuclear** movements of the molecular systems

ZPVC - Zero Point Vibrational Correction^{a b}

^aP-O. Astrand et al., J. Chem. Phys., 102, 3534, (1995).

^bK. Ruud et al., J. Chem. Phys., 112, 2668, (2000).

- Dalton 2.0/2016 - **HF/SCF**, **MCSCF** (CAS/RAS) e **DFT**
- Variational expansion point - effective geometry (vibrationally averaged)
- **2^a order** Perturbation Theory - feasible to polyatomics



Vibrationally Averaged Geometry and Dipole Moments

Effective Geometry

$$r_{ef} = \langle r_i \rangle = r_{e,i} - \frac{1}{4\omega_i} \sum_{j=1}^{3N-6} \frac{V_{ij}^{(3)}}{\omega_j} \quad (8)$$

Perturbative expansion of the vibrational wavefunction $|\Psi\rangle_{vib}$

Vibrationally Averaged Dipole Moments

$$\langle \mu \rangle = \mu_{ef} + \frac{1}{4} \sum_i \frac{\mu_{ef,ii}^{(2)}}{4\eta\omega_{ef}} \quad (9)$$

η is the reduced mass of the system



FNMC/ZPVC Dipole Moments (Dalton 2016 Program)

FNMC/ZPVC predicts highly accurate **dipole moments** to

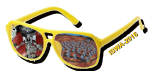
- HD , HT , CH_3CD_3 , $^{13}CH_3CD_3$, $CH_3^{13}CD_3$, CH_2CD_2 (*asym.*), $CHDCHD$ (*cis*) - Originally **apolar** systems.⁸
- C_3H_8 , $CH_3CD_2CH_3$, $CD_3CH_2CD_3$, $CHD_2CH_2CHD_2$, CH_3CCH , CH_3CCD , CD_3CCH , CD_3CCD , H_2O , $[HDO]$, D_2O - Originally **polar** systems.⁹

Two important findings

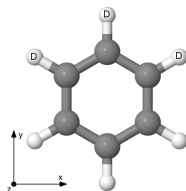
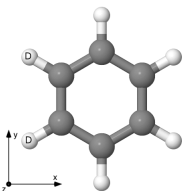
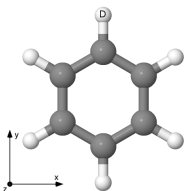
- The isotopic substitution **trend**.
- **Accuracy** relative to experimental data - **Benchmark!**.

⁸Arapiraca et al., J. Chem. Phys. 135, 244313 (2011)

⁹Arapiraca and Mohallem., Chem. Phys. Let., 609, 123 (2014)



FNMC/ZPVC Dipole Moments (Dalton 2016 Program)

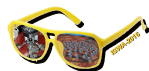


B3LYP/FNMC/ZPVC/6-31G**	Dipole	Moments	
System	$\langle \mu \rangle_z^{10}$	Experiment ¹¹	$\Delta(\%)$
C_6H_5D	0.0078	0.0081 ^a	3.7%
$C_6H_4D_2$	0.0134	-	-
$C_6H_3D_3$	0.0154	-	-

Vibrationally averaged isotopic dipole moments of benzene isotopologues in debye units at B3LYP/FNMC/ZPVC/6-31G** level

¹⁰Arapiraca et al., J. Chem. Phys. 144, 144301 (2016)

¹¹M. Oldani and A. Bauder, Chem. Phys. Let., 108, 7 (1984)



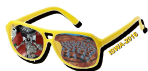
ZPVC Rotational Constants (Gaussian 09 program)

	B3LYP/ZPVC/aug-cc-pVTZ	Experiment	$\Delta(\%)$
$C_6H_5D^{12}$			
A	5692.143	5689.144(6)	0.053%
B	5326.934	5323.934(6)	0.056%
C	2750.913	2749.674(6)	0.045%
$C_6H_4D_2^{13}$			
A	5501.102	5498.062	0.062%
B	5167.642	5164.242	0.066%
C	2663.735	2662.496	0.046%
$C_6H_3D_3$			
A	5217.7295	-	-
B	5201.4588	-	-
C	2604.7907	-	-

Rotational constants of benzene isotopologues in MHz units.

¹²M. Oldani et al., J. Mol. Struc., 1909, 190 (1983)

¹³M. Oldani and A. Bauder, Chem. Phys. Let., 108, 7 (1984)

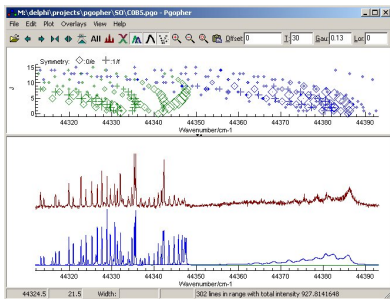


Simulating Rotational Spectra (PGOPHER Program)

Programa PGOPHER

To resolve the molecular **rotational structure**^a

^aPGOPHER, a Program for Simulating Rotational Structure, C. M. Western, University of Bristol, <http://pgopher.chm.bris.ac.uk>



Simulating Rotational Spectra (PGOPHER Program)

Polyatomic Molecule - Asymmetric Top

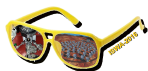
Inertia moments and angular moment projections

$$I_A \leq I_B \leq I_C \quad (10)$$

$$A \leq B \leq C \quad (11)$$

$$\hat{H} = A\hat{J}_a^2 + B\hat{J}_b^2 + C\hat{J}_c^2 \quad (12)$$

$$\kappa = \frac{2B - A - C}{A - C} \quad (13)$$



Simulating Rotational Spectra (PGOPHER Program)

Einstein's Coefficients

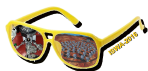
Dipole moment approximation - Emission and Absorption

$$N_0 B_{1 \leftarrow 0} \rho_\nu = A_{1 \rightarrow 0} N_1 + B_{1 \rightarrow 0} \rho_\nu N_1 \quad (14)$$

$$B_{1 \leftarrow 0} = B_{1 \rightarrow 0} \quad (15)$$

$$A_{1 \rightarrow 0} = \frac{8\pi h \nu_{10}^3}{c^3} B_{1 \leftarrow 0} \quad (16)$$

$$A_{1 \rightarrow 0} = \frac{16\pi^3 \nu^3}{3\epsilon_0 h c^3} \mu_{10}^2 \quad (17)$$



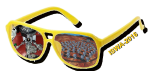
Simulating Rotational Spectra (PGOPHER Program)

Computing state populations

Emission and Absorption Intensities

$$I(T)_{emiss} = \frac{64\pi^4 \nu^3 S_g \mu_g^2 (e^{-E''/kT} - e^{-E'/kT})}{3hc Q} \quad (14)$$

$$I(T)_{abs} = \frac{8\pi^3 \nu S_g \mu_g^2 (e^{-E''/kT} - e^{-E'/kT})}{3hc Q} \quad (15)$$



Multiplicative Mean Correction Factor - MCF

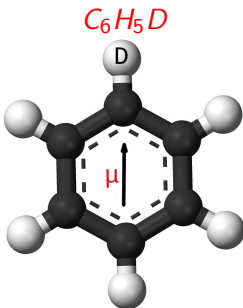
$$CF_i = \frac{V_i^{\text{exp}}}{V_i^{\text{teo}}} \quad (16)$$

$$MCF = \frac{1}{N} \sum_{i=1}^N FC_i \quad (17)$$

$$V_{C_i}^{\text{teo}} = V_i^{\text{teo}} MCF \quad (18)$$



Benzene's Isotopologues Spectra



$$C_{2v} - \kappa = 0.75$$

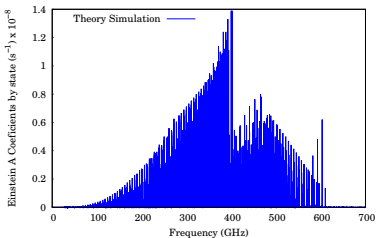
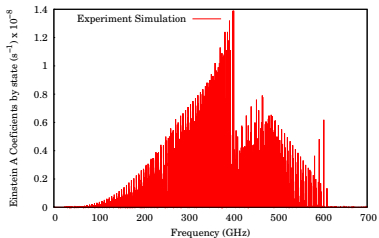
Watson Asymmetric Reduction III'



Benzene's Isotopologues Spectra

C_6H_5D Spontaneous Emission Rates by State

29253 lines: $0 \leq J \leq 40 - 0 \leq K \leq 31$



MAPE error without MCF correction

Frequency MAPE error: 0.08% - Intensity MAPE error: 7.51%

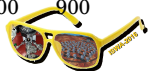
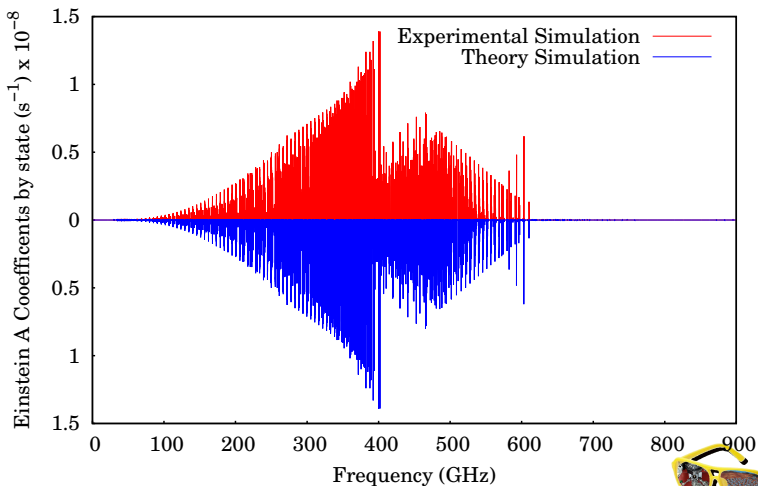
MAPE error with MCF correction

Frequency MAPE error: 0.03% - Intensity MAPE error: 0.44%



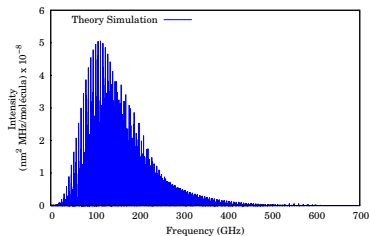
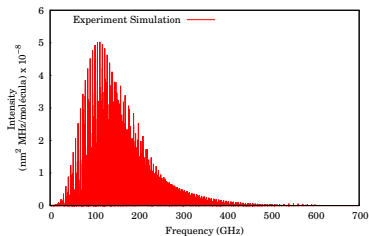
Benzene's Isotopologues Spectra

C_6H_5D Spontaneous Emission Rates
 $0 \leq J \leq 40$ and $0 \leq K \leq 31$



Benzene's Isotopologues Spectra

C_6H_5D Absorption Spectra at 35 K
29253 lines: $0 \leq J \leq 40 - 0 \leq K \leq 31$

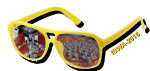


MAPE error without MCF correction

Intensity MAPE error: 7.78%

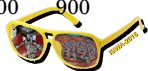
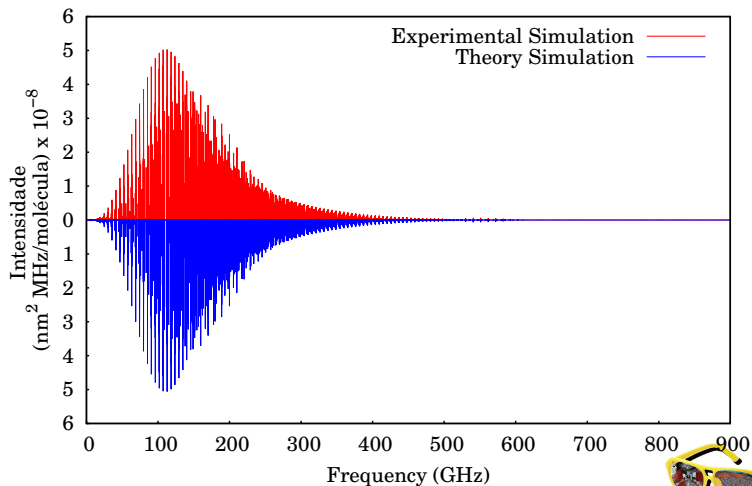
MAPE error with MCF correction

Intensity MAPE error: 0.45%

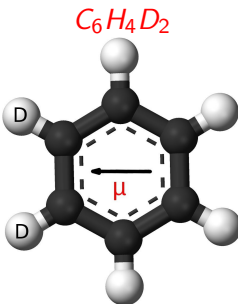


Benzene's Isotopologues Spectra

C_6H_5D Absorption Spectra at 35 K
 $0 \leq J \leq 40$ and $0 \leq K \leq 31$

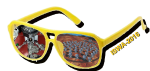


Benzene's Isotopologues Spectra



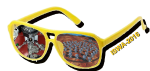
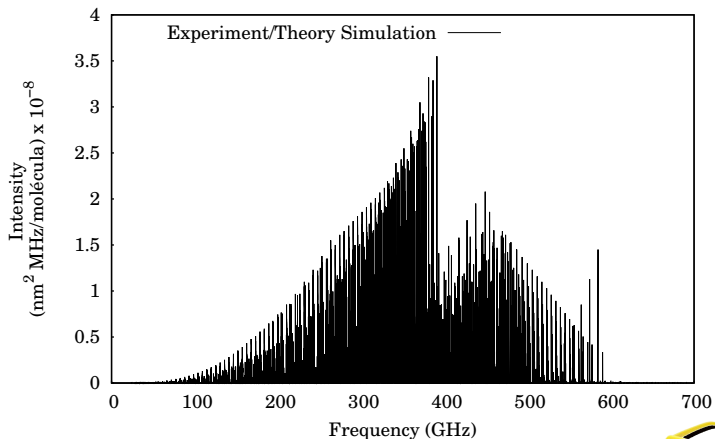
$$C_{2v} - \kappa = 0.76$$

Watson Asymmetric Reduction III^r



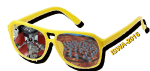
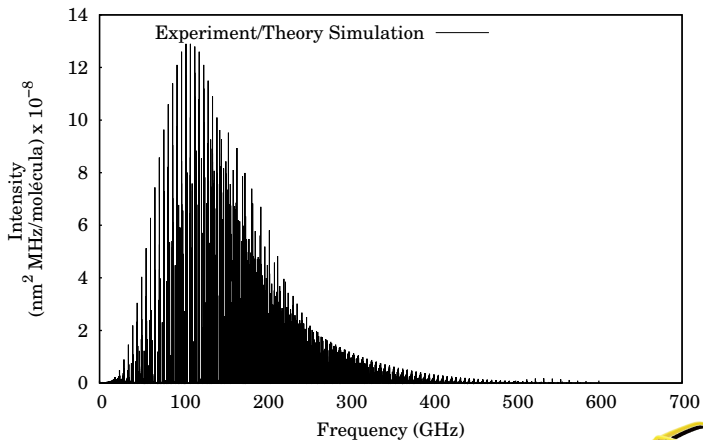
Benzene's Isotopologues Spectra

$C_6H_4D_2$ Spontaneous Emission Rates by State
 $0 \leq J \leq 40$ e $0 \leq K \leq 31$



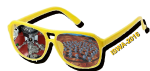
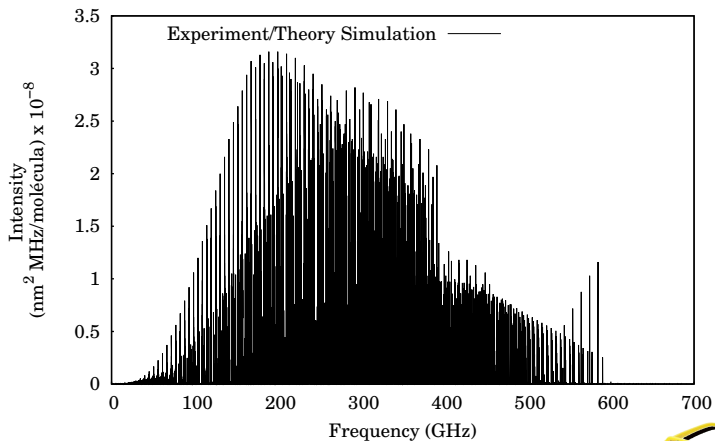
Benzene's Isotopologues Spectra

$C_6H_4D_2$ Absorption Spectra at 35 K
 $0 \leq J \leq 40$ and $0 \leq K \leq 31$

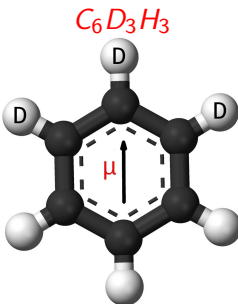


Benzene's Isotopologues Spectra

$C_6H_4D_2$ Absorption Spectra at 150 K
 $0 \leq J \leq 40$ and $0 \leq K \leq 31$

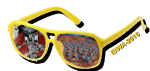


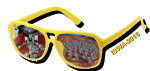
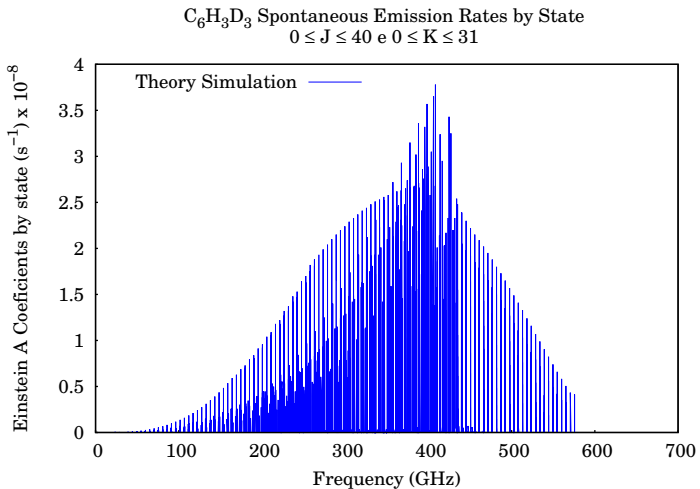
Benzene's Isotopologues Spectra



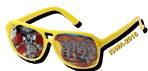
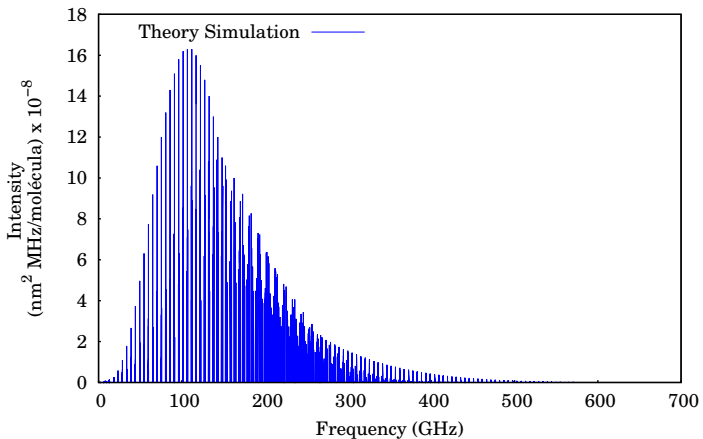
$$C_{2v} - \kappa = 0.99$$

Watson Asymmetric Reduction III^r

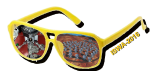
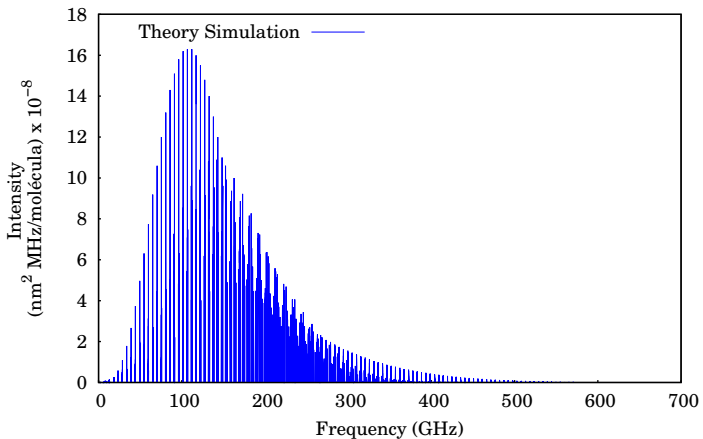




$C_6H_4D_2$ Absorption Spectra at 35 K
 $0 \leq J \leq 40$ and $0 \leq K \leq 31$

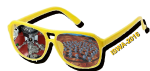


$C_6H_3D_3$ Absorption Spectra at 35 K
 $0 \leq J \leq 40$ and $0 \leq K \leq 31$



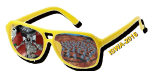
Conclusions

- Theoretical description of isotopic effects in molecules.
- Efficient Electronic and vibrational analysis.
- *Ab Initio* rotational spectra to deuterated molecules.
- Several proposals of experiments.

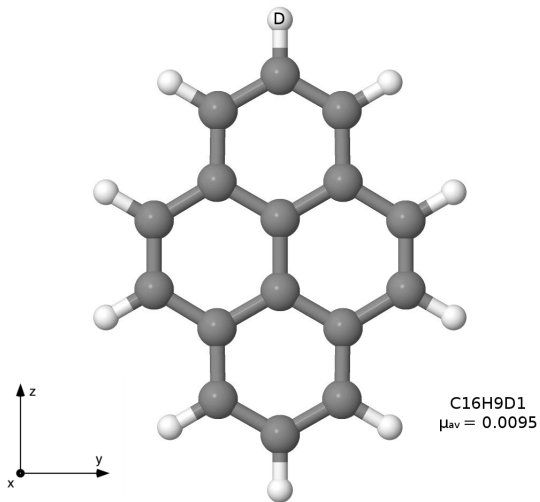


Perspectives

- Improving the error analysis.
- Theoretical support to detection of molecules.
- Line survey of SGR-B2 unidentified lines (Sergio Pilling).
- Column densities and abundances (RADEX...)(Edgar Mendoza).
- Deuterated benzene's formation process (João Bosco da Silva).
- PAH rotational spectra.



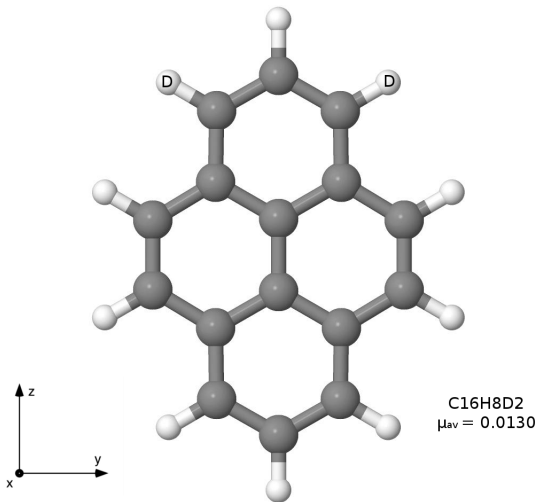
Coming Soon



Jmol



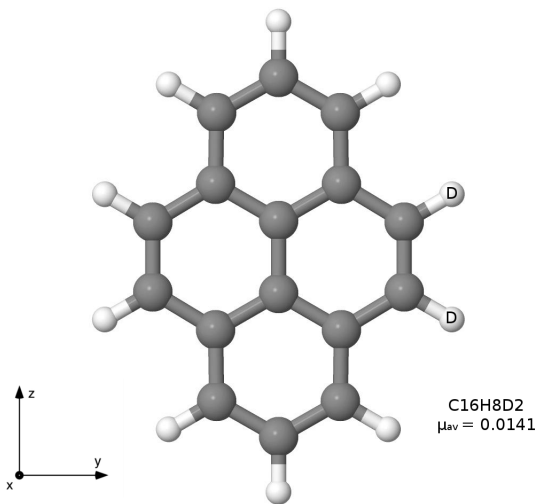
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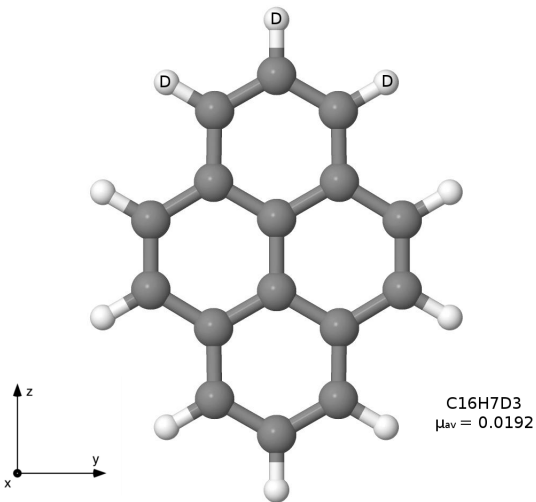
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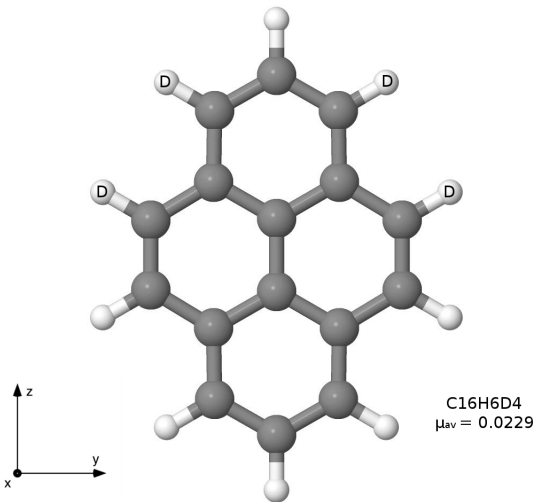
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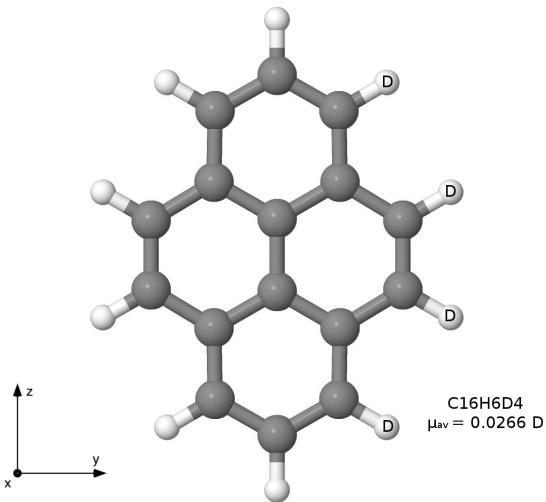
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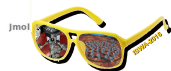
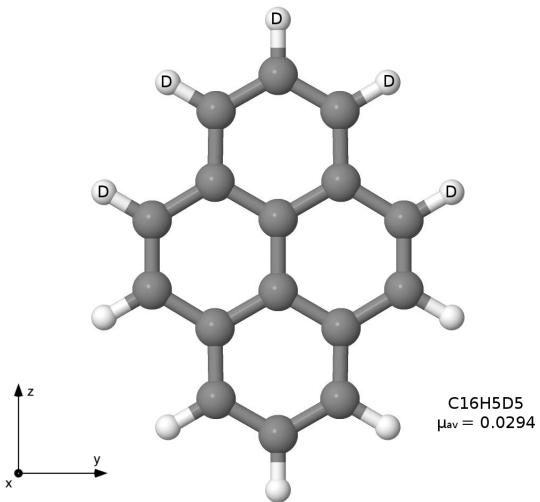
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Acknowledgment

