

Guanine formation mechanism from observed interstellar chemical species

João Bosco Paraiso da Silva and Ana Paula M. de Araujo

Departamento de Química Fundamental – UFPE - Brazil

paraiso@ufpe.br

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Summary

Facts about (bio)molecules in the ISM

Methodological background

Learnings from Adenine mechanism of Synthesis

Results on the Guanine Synthesis

Conclusions

~ 150 molecules were observed in the ISM

Table 1 Interstellar molecules^a (cumulenes and heterocumulenes are shown in bold letters)

Diatomic

AlCl, AlF, AlO, CC, CH, CH⁺, CF⁺, CN, CN⁻, CO, CO⁺, CP, CS, CSi, FeO, HCl, HF, H₂, HN, HO, HS, KCl, LiF, MgH⁺, N₂, NO, NP, NS, NSi, NaCl, NaI, O₂, OP, OS, OS⁺, OSi, SH, SSi.

Triatomic

AlNC, AlOH, C₃, **C₂H**, CH₂, **C₂O**, CO₂, **C₂P**, **C₂S**, *cyclo*-C₂Si, HCN, HCO, HCO⁺, HCP, HCS, HCS⁺, HDO, H₃⁺, HNC, HN₂⁺, HNO, H₂D⁺, HD₂⁺, H₂N, H₂O, H₂S, HOC⁺, KCN, MgCN, MgNC, NaCN, NaOH, N₂O, OCN⁻, OCS, O₃, SO₂, SiCN, SiNC.

4-atomic

CH₂O, CH₂S, C₂H₂, **C₂HN**, CH₂N, C₃H, *cyclo*C₃H, C₃N, C₃O, C₃S, C₃Si, CH₃, C₄, HCNH⁺, HCNO, HNCO, HNCN, HSCN, HOCO⁺, H₃O⁺, H₂O₂, NH₃.

5-atomic

CH₄, CH₂CN, CH₂CO, CH₂NC, CH₂NH, CH₂OH⁺, C₃H₂, *cyclo*C₃H₂, HC₂CN, HC₂NC, C₃NH, HC₃N, C₄H, C₄H⁻, C₄Si, C₅, HCO₂H, HCOCN, NH₂CN, SiH₄.

6-atomic

C₂H₄, CH₂CHO, CH₂CNH, *cyclo*C₃H₂O, CH₃CN, CH₃NC, CH₃OH, CH₃SH, C₃H₂N, C₄H₂, C₄HN, C₅H, C₅N⁻, HC₂CNH⁺, HCONH₂, NH₂CHO.

7-atomic

*cyclo*C₂H₄O, CH₂CHCN, CH₂CHOH, CH₃CHO, CH₃NH₂, CH₃C₂H, HC₄CN, C₆H, C₆H⁻.

8-atomic

CH₂OHCHO, CH₂CHCHO, CH₂CCHCN, CH₃CO₂H, CH₃C₂CN, C₆H₂, C₇H, HCO₂CH₃, NH₂CONH₂, NH₂CH₂CN.

9-atomic

CH₃CH₂CN, CH₃CH₂OH, CH₃CHCH₂, CH₃CONH₂, CH₃OCH₃, CH₃C₄H, C₇HN, C₈H, C₈H⁻.

10 or more atoms

CH₃CH₂CHO, CH₃COCH₃, CH₃C₄CN, HOCH₂CH₂OH, CH₃C₆H, HC₈CN, HCO₂C₂H₅, *n*C₃H₇CN, C₆H₆, HOCH₂COCH₂OH, HC₁₀CN, HC₁₁N, C₁₄H₁₀, C₆₀, C₇₀.

Deuterated molecules

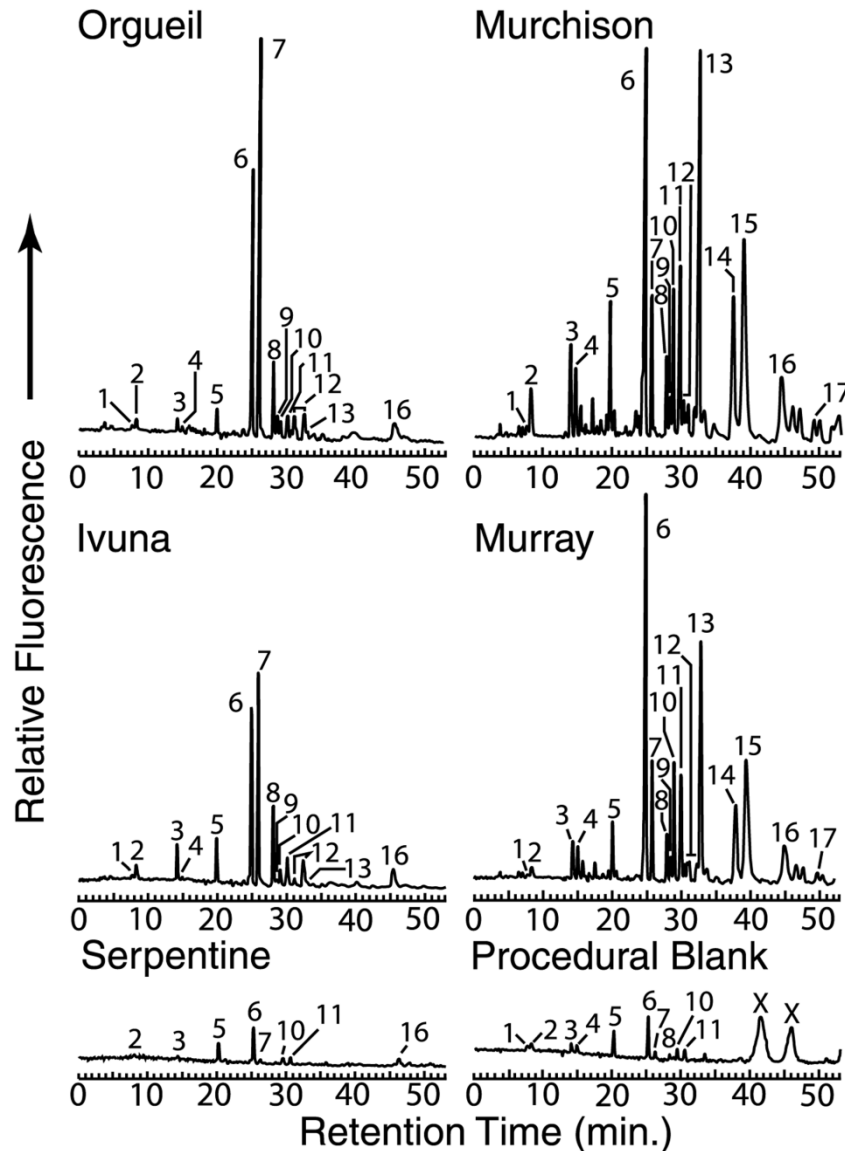
HD, H₂D⁺, HD₂⁺, HDO, D₂O, DCN, DCO, DNC, N₂D⁺, NH₂D, NHD₂, ND₃, HDCO, D₂CO, CH₂DC₂H, CH₃C₂D.

Molecules reported but not confirmed

HOCN, NH₂CH₂CO₂H, CO(CH₂OH)₂, C₂H₅COCH₃, C₁₀H₈⁺, SiH, PH₃.

^aThis list was compiled from data listed in: <http://www.astro.uni-koeln.de/cdms/molecules> http://www.astrochymist.org/astrochymist_ism.html http://en.wikipedia.org/wiki/List_of_molecules_in_interstellar_space, also ref. 43.

Chromatogram Samples extracted from Meteors



- 1 D-Aspartic Acid
- 2 L-Aspartic Acid
- 3 L-Glutamic Acid
- 4 D-Glutamic Acid
- 5 D,L-Serine
- 6 Glycine
- 7 β -Alanine
- 8 γ -Amino-*n*-butyric Acid (g-ABA)
- 9 D,L-b-Aminoisobutyric Acid (b-AIB)
- 10 D-Alanine
- 11 L-Alanine
- 12 D,L- β -Amino-*n*-butyric Acid (b-ABA)
- 13 α -Aminoisobutyric Acid (AIB)
- 14 D,L- α -Amino-*n*-butyric Acid (a-ABA)
- 15 D,L-Isovaline
- 16 L-Valine
- 17 D-Valine
- X: unknown

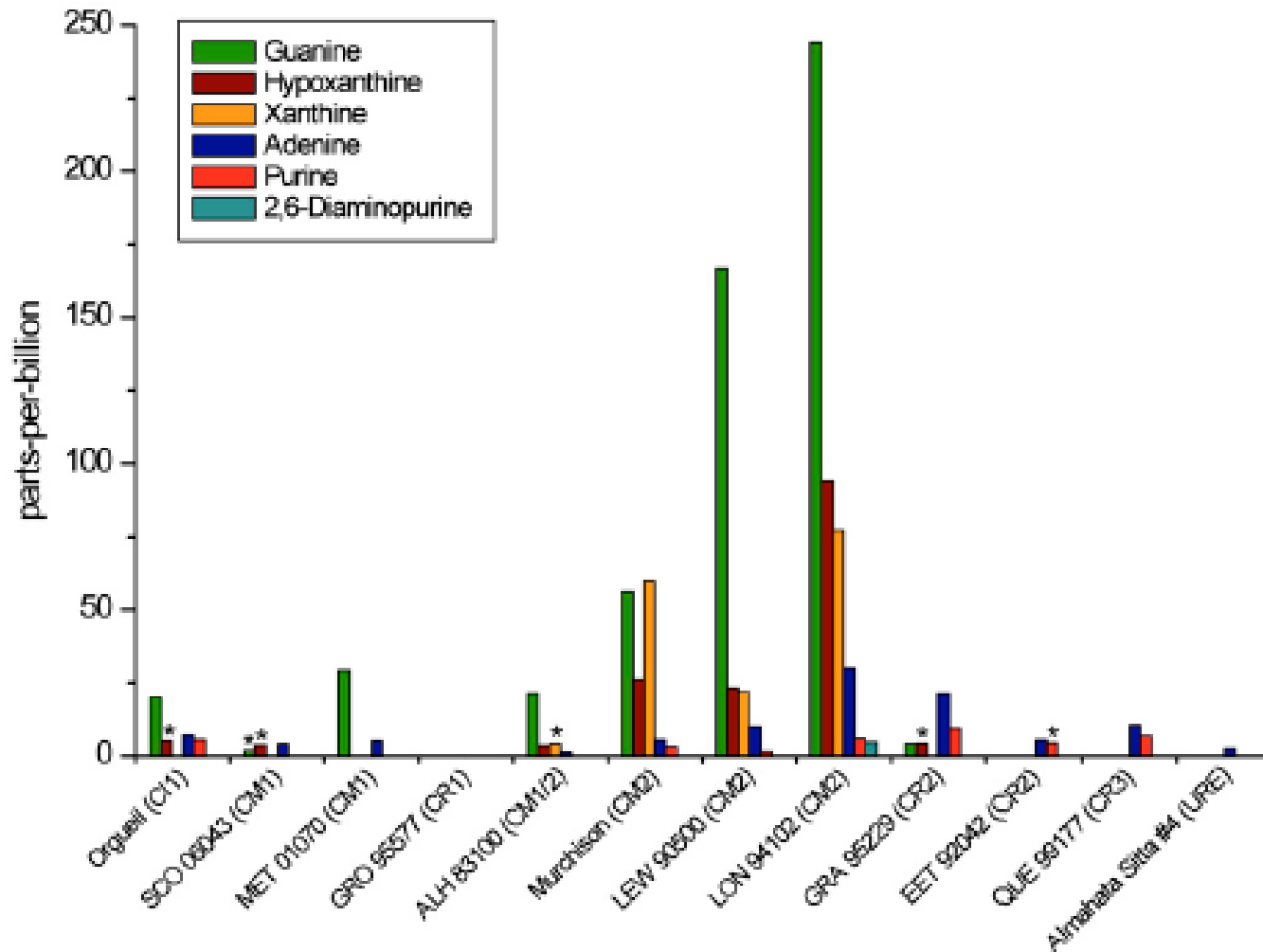


Fig. 1. Distribution of guanine, hypoxanthine, xanthine, adenine, purine, and 2,6-diaminopurine in 11 carbonaceous chondrites and one ureilite.

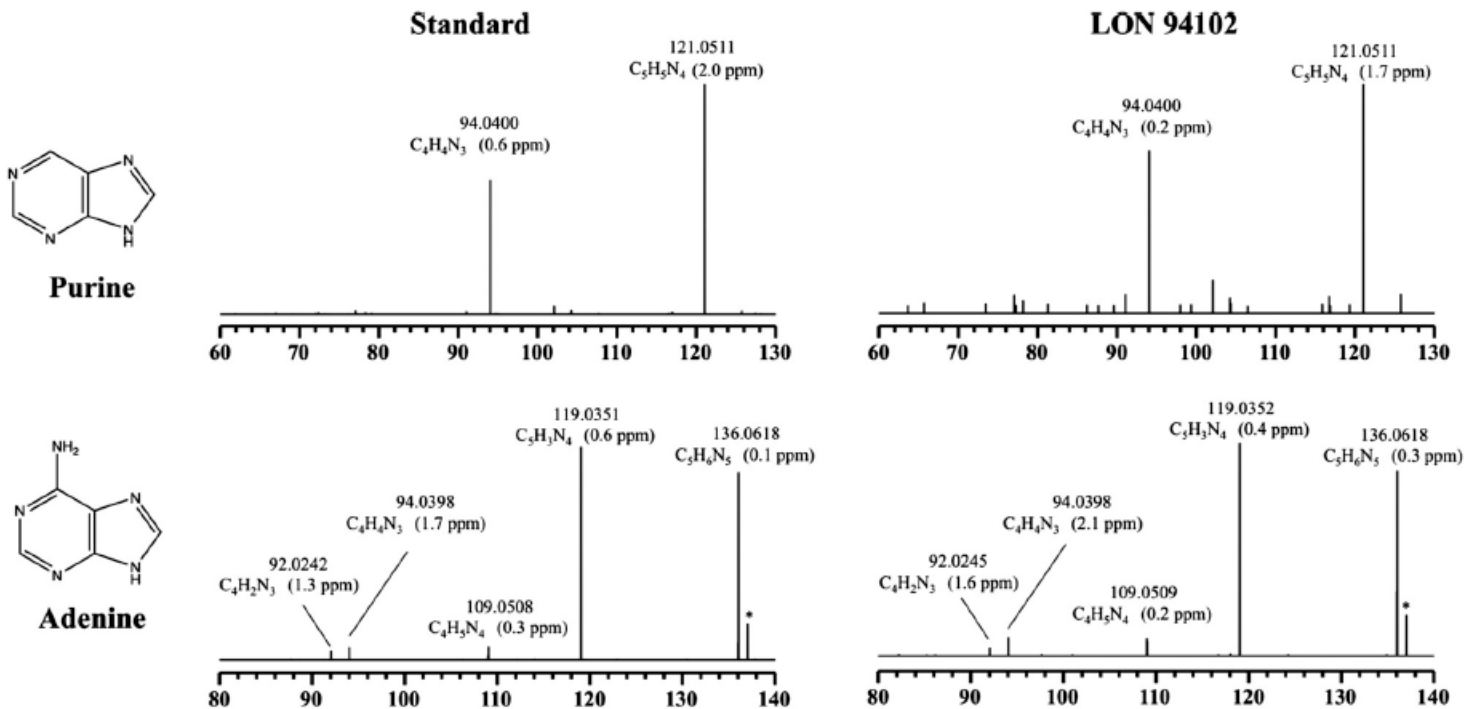


Fig. 2. Mass-selected fragmentation spectra of reference standards (left spectra) and compounds found in the meteorite LON 94102 (right spectra) measured on an LTQ Orbitrap XL hybrid mass spectrometer using an HCD (higher energy collision dissociation) setting of 90 to 100%. Purine, adenine, 2,6-diaminopurine, and 6,8-diaminopurine were identified using accurate mass measurements on the parent mass and multiple fragment masses and chromatographic retention time. Mass accuracy of less than 5 ppm allows for the unambiguous assignment of elemental formulae. The * represents inferences in the fragmentation spectra that are present in both the meteorite and reference standard spectra.

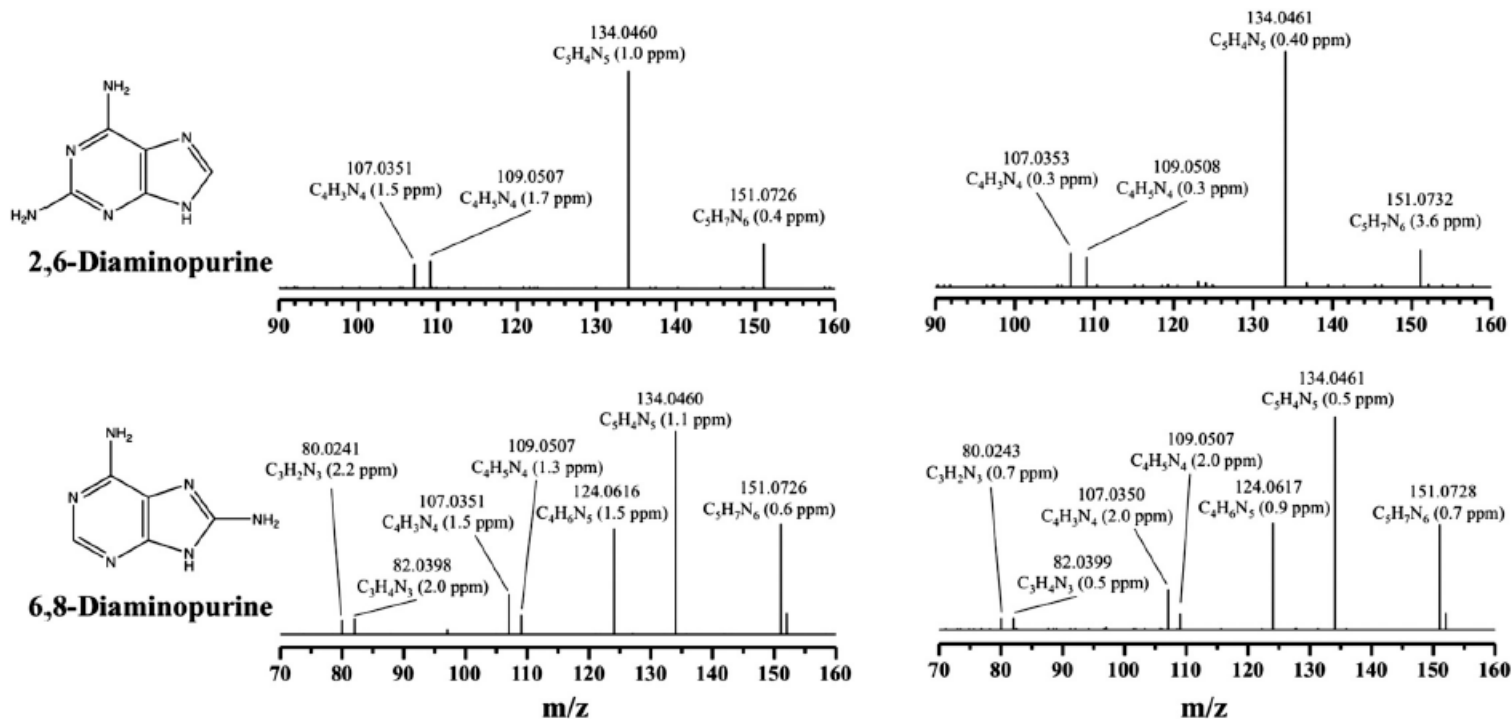
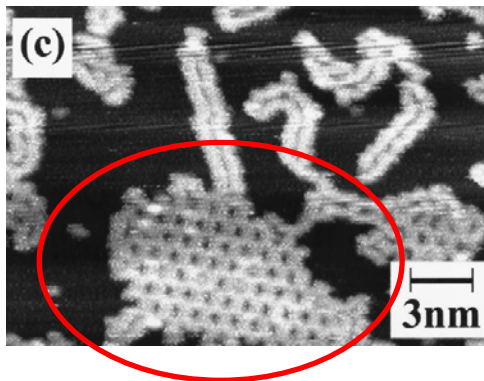
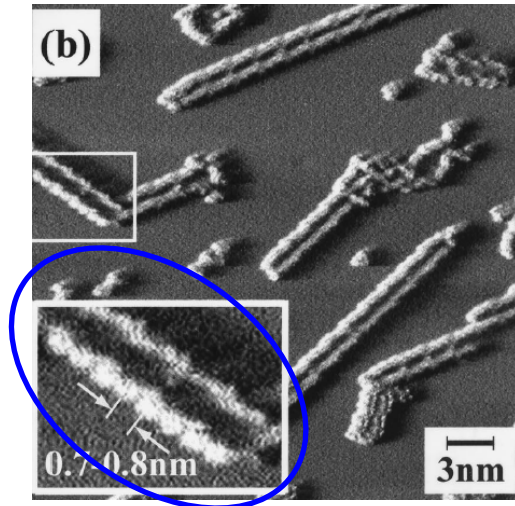
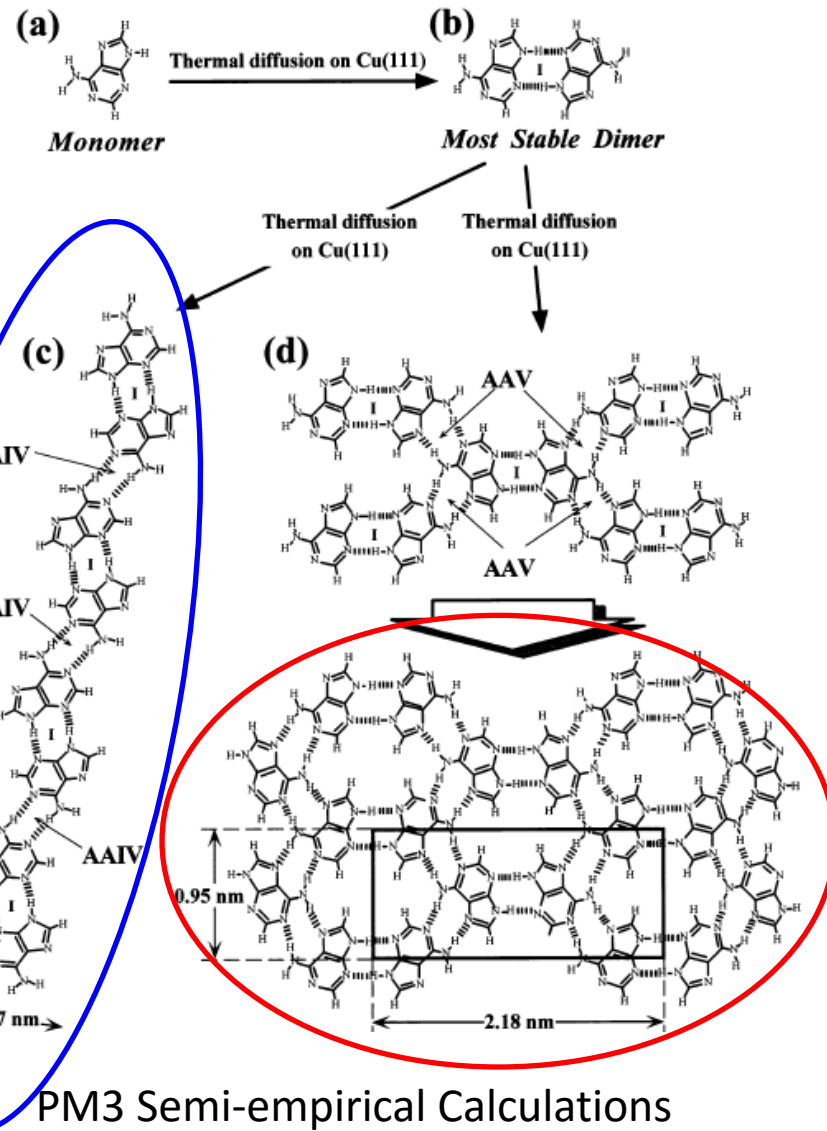


Fig. 2. Mass-selected fragmentation spectra of reference standards (left spectra) and compounds found in the meteorite LON 94102 (right spectra) measured on an LTQ Orbitrap XL hybrid mass spectrometer using an HCD (higher energy collision dissociation) setting of 90 to 100%. Purine, adenine, 2,6-diaminopurine, and 6,8-diaminopurine were identified using accurate mass measurements on the parent mass and multiple fragment masses and chromatographic retention time. Mass accuracy of less than 5 ppm allows for the unambiguous assignment of elemental formulae. The * represents inferences in the fragmentation spectra that are present in both the meteorite and reference standard spectra.

Low-dimensional of super hydrogen-bond complexes of adenine on Cu(111) surfaces

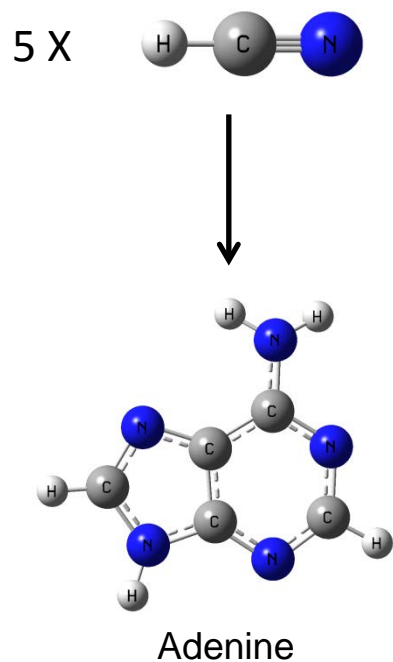


Scanning Tunneling Microscope (STM)

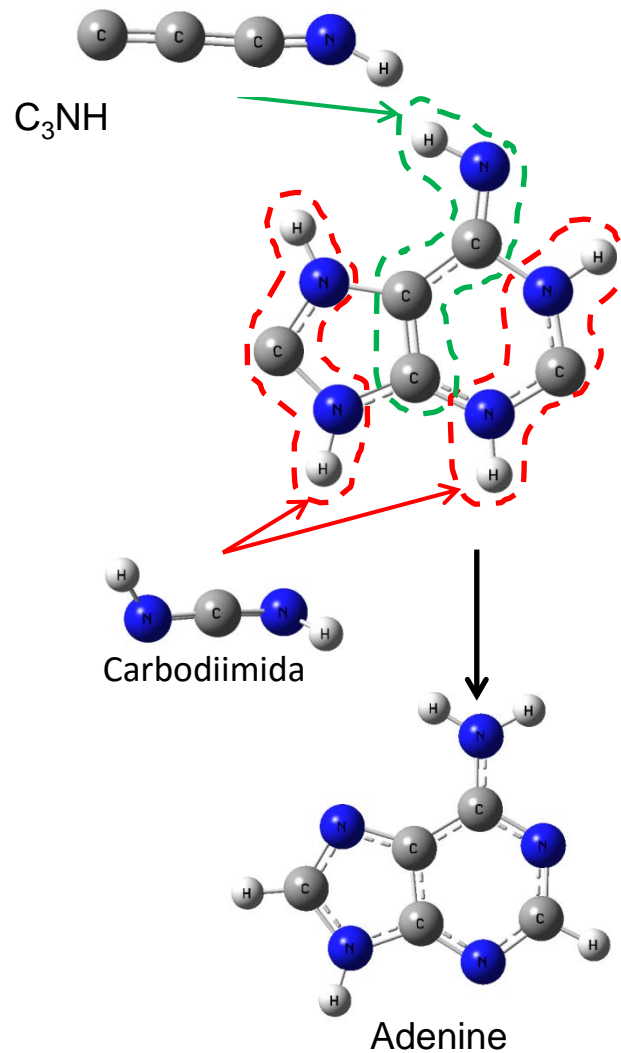


Mechanisms for the Synthesis of Adenine

D. Roy, K. Najafian, P. von R. Schleyer
PNAS 104, 17272 (2007).



J.B.P. da Silva, E. C. de Aguiar, K. M. Merz
J. Chem. Phys. A 118, 3637 (2014).



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B3LYP/6-311+G(d,p)

Rate determinant step \rightarrow 71 kcal mol⁻¹

First it forms the 5-member ring,
then he 6-member ring

Involves 14 steps

$\Delta G_R = -53.7$ kcal mol⁻¹

J.B.P. da Silva, E. C. de Aguiar, K. M. Merz
J. Chem. Phys. A 118, 3637 (2014).

MP2/6-311++G(d,p)

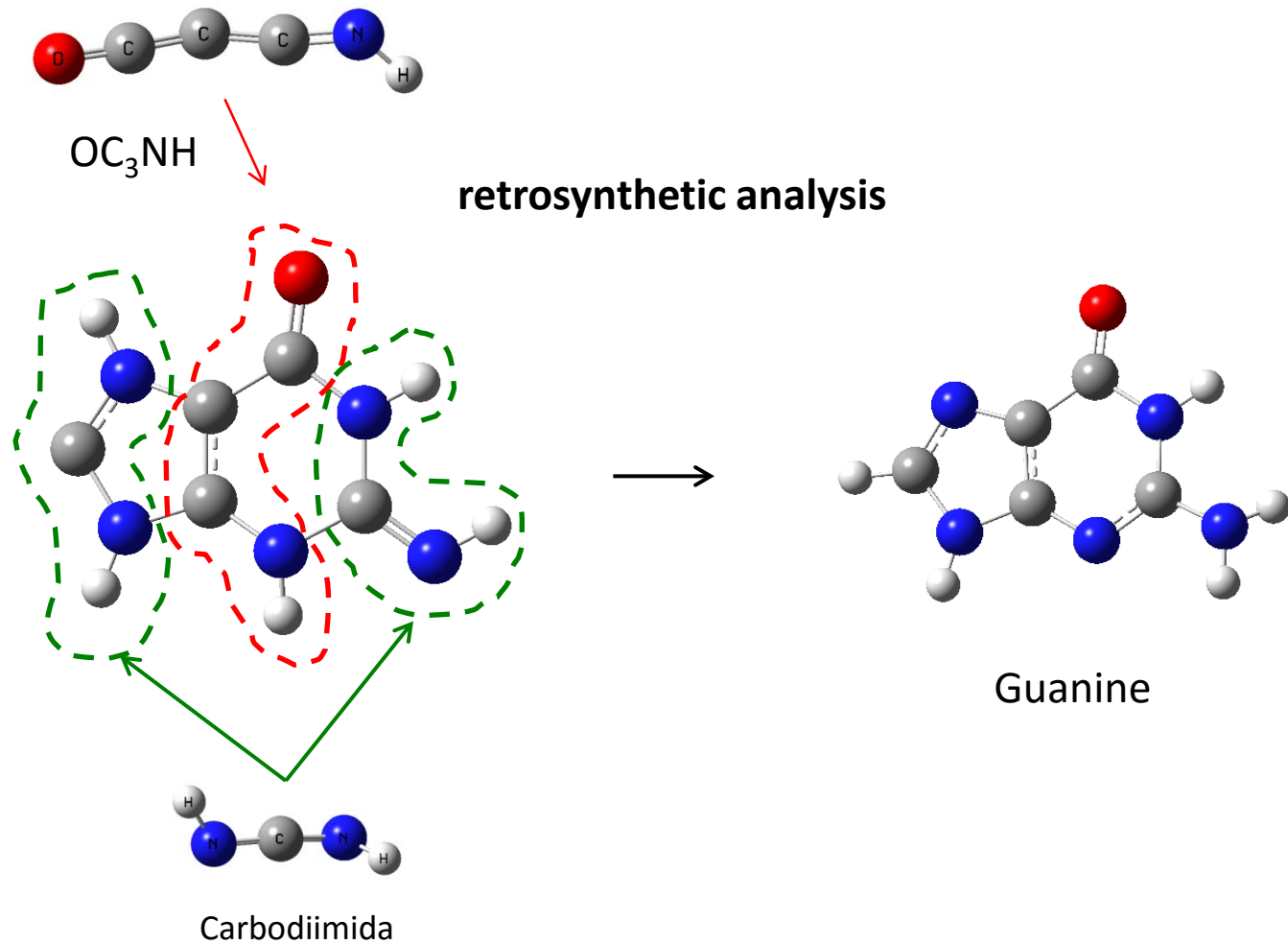
Rate determinant step \rightarrow 23 kcal mol⁻¹

First it forms the 6-member ring,
then he 5-member ring

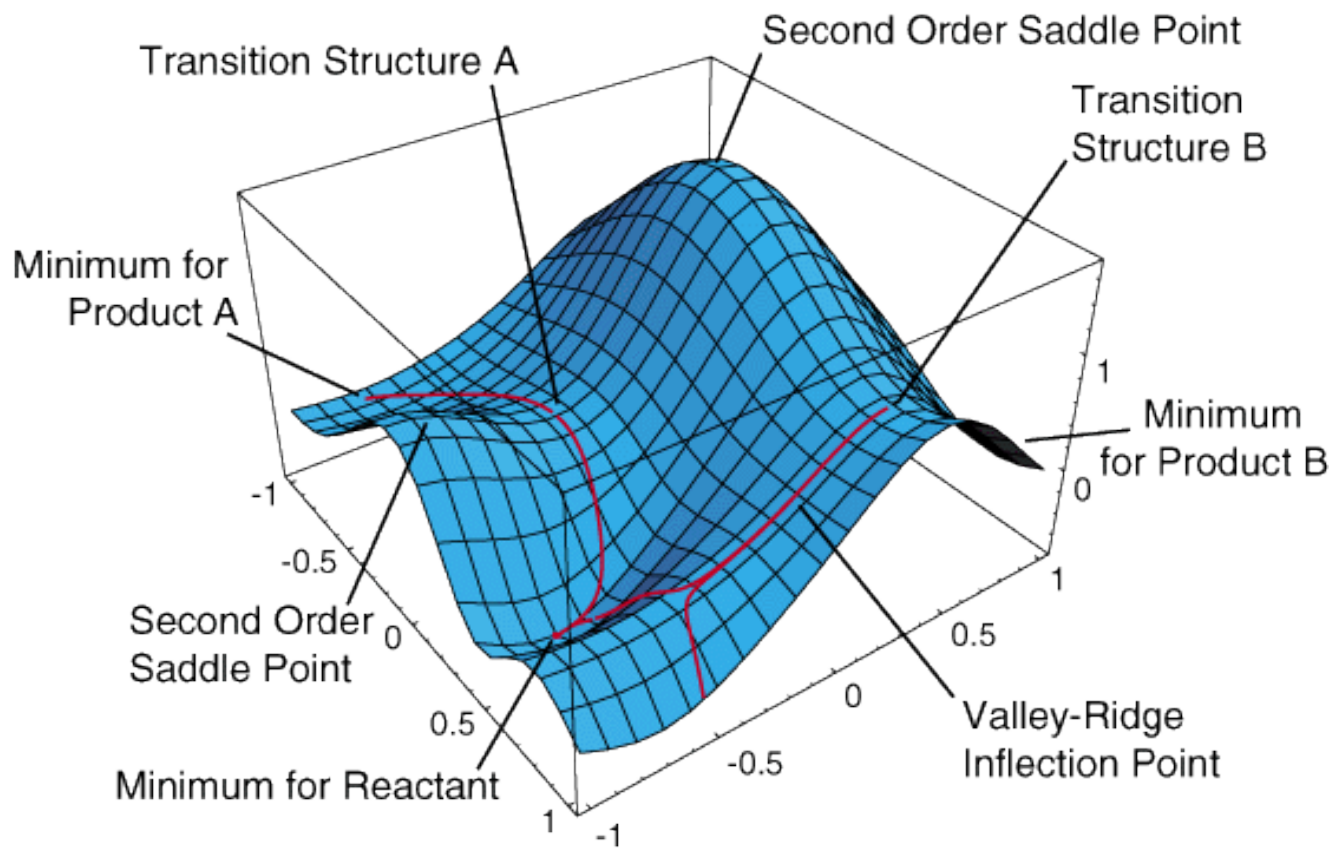
Involves 6 steps

$\Delta G_R = -142.2$ kcal mol⁻¹

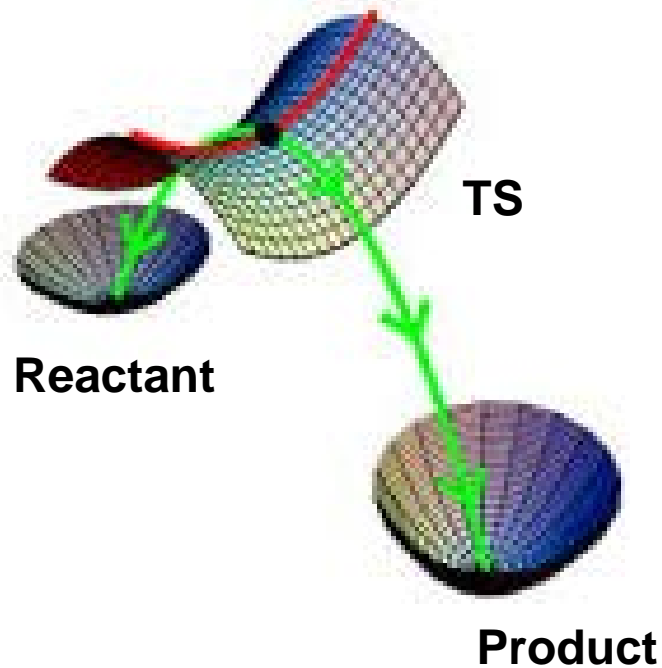
Mechanism for the Synthesis of Guanine



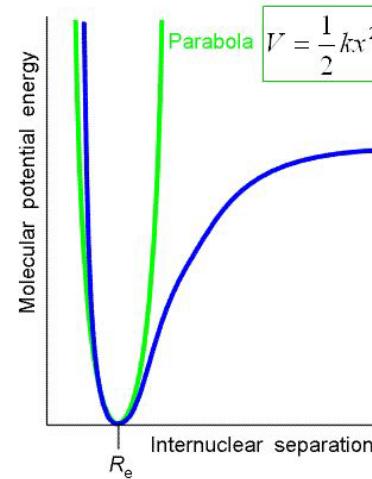
Potential Energy Surface (PES)



Potential Energy Surface (PES)



Harmonic approximation



$$f(x) = \frac{1}{2\pi} \left(\frac{k}{\mu} \right)^{1/2}$$

$$k = \left(\frac{\partial^2 V}{\partial x^2} \right)_0$$

For real reactants and products $k > 0 \rightarrow f(x)$ is a real number

For saddle point $k < 0 \rightarrow f(x)$ is a imaginary number

Methodological background

Target: Solve the time independent Schrödinger equation

$$H\Psi = E\Psi$$

$$H = \sum_{i=1}^n \left(-\frac{\hbar^2}{2m_i} \nabla_i^2 \right) + \sum_{A=1}^N \left(-\frac{\hbar^2}{2M_A} \nabla_A^2 \right) + \sum_{i=1}^n \sum_{A=1}^N \left(-\frac{Ze^2}{r_{iA}} \right) + \sum_{i=1}^n \sum_{j>1}^n \frac{e^2}{r_{ij}} + \sum_{A=1}^N \sum_{B>A}^N \frac{Z_A Z_B e^2}{R_{AB}}$$

$K_{\text{electrons}}$

K_{nuclei}

$V_{\text{el.-nucl.}}$

$V_{\text{el.-el..}}$

$V_{\text{nucl.-nucl.}}$

Ψ – wave function; E – total energy of the system

$$H = E + KT$$

$$G = H - TS$$

Computation

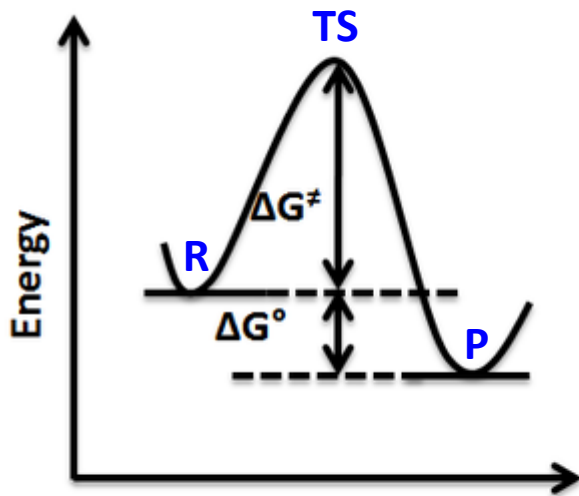
G09, MP2/6-311++G(d,p), gas phase, default internal criteria of convergence,

energy in kcal mol⁻¹.

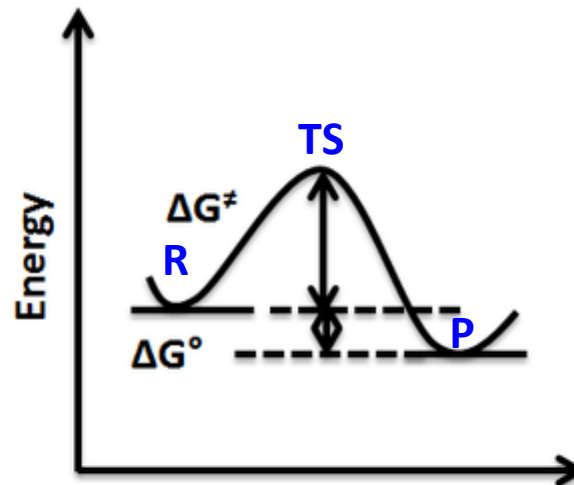
T = 10 K

P = 1 bar

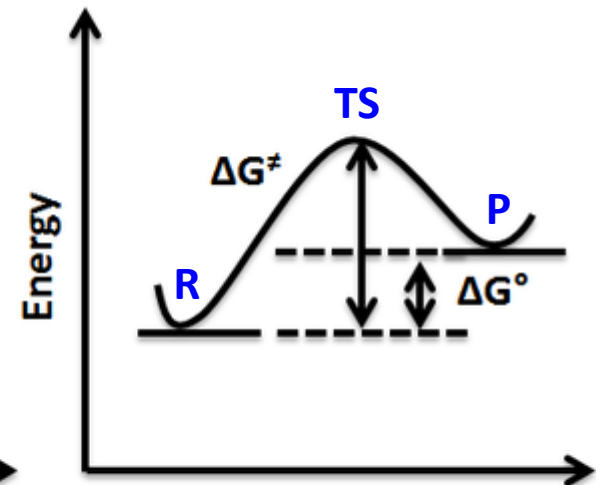
Methodological background



Large ΔG^\ddagger , $\Delta G^\circ < 0$
Slow and favorable

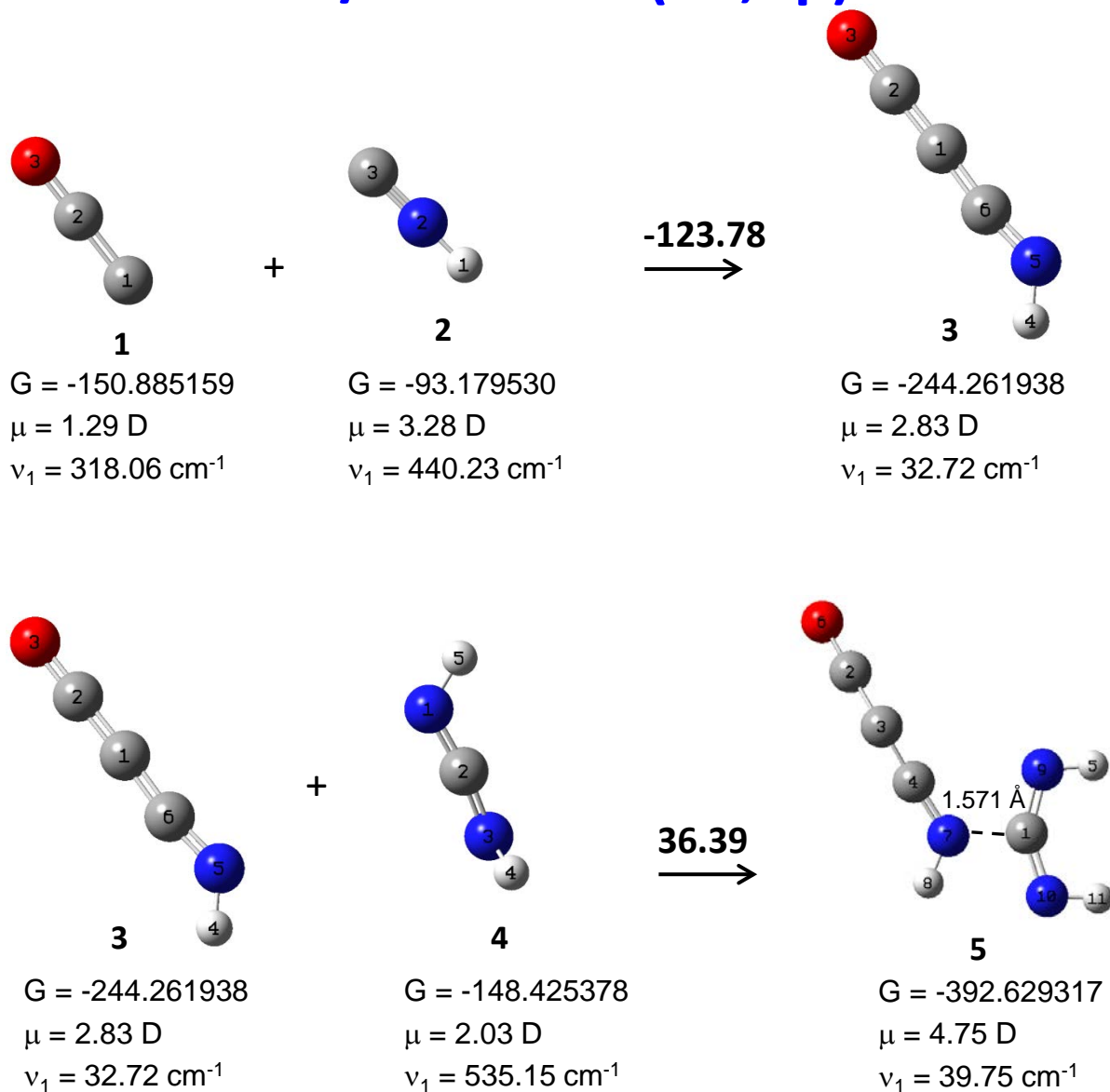


Small ΔG^\ddagger , $\Delta G^\circ < 0$
Fast and favorable

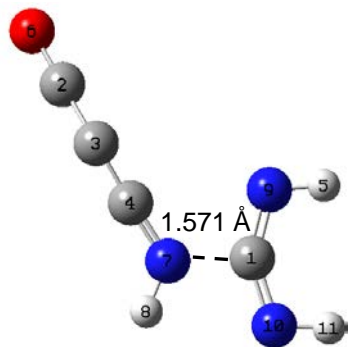


Small ΔG^\ddagger , $\Delta G^\circ > 0$
Fast and unfavorable

MP2/6-311++G(2d,2p)



MP2/6-311++G(2d,2p)



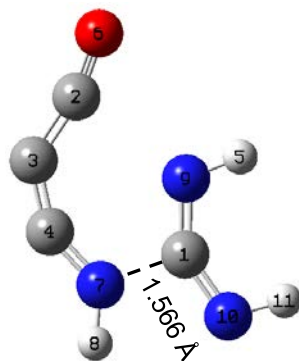
5

$G = -392.629317$

$\mu = 4.75 \text{ D}$

$\nu_1 = 39.75 \text{ cm}^{-1}$

8.92
→



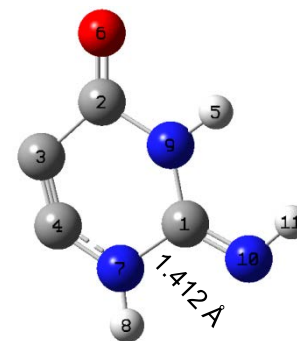
TS_{5,6}

$G = -392.615104$

$\mu = 2.69 \text{ D}$

$\nu_1 = -276.05 \text{ cm}^{-1}$

-26.61
→

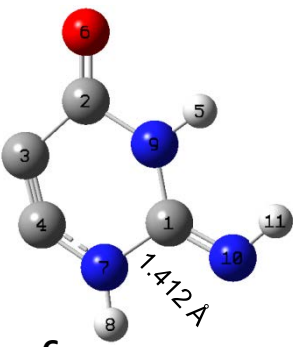


6

$G = -392.657509$

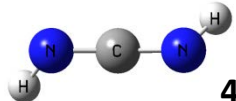
$\mu = 2.43 \text{ D}$

$\nu_1 = 26.09 \text{ cm}^{-1}$



6

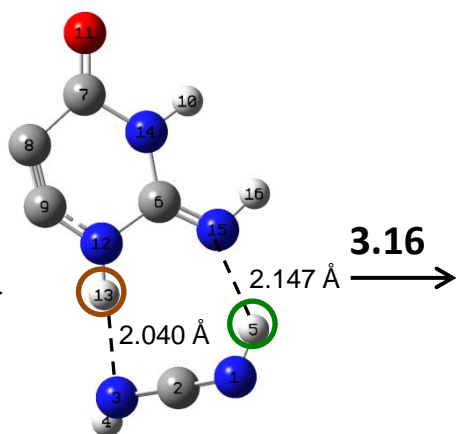
+



4

$G = -541.082887$

-8.45
→



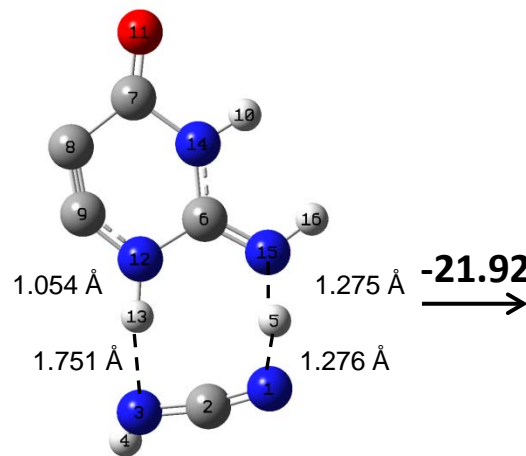
7

$G = -541.096366$

$\mu = 2.93 \text{ D}$

$\nu_1 = 27.16 \text{ cm}^{-1}$

3.16
→



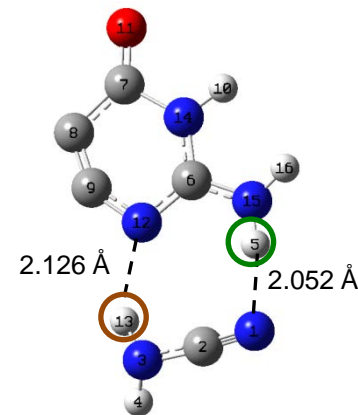
TS_{7,8}

$G = -541.091327$

$\mu = 2.78 \text{ D}$

$\nu_1 = -1069.36 \text{ cm}^{-1}$

-21.92
→



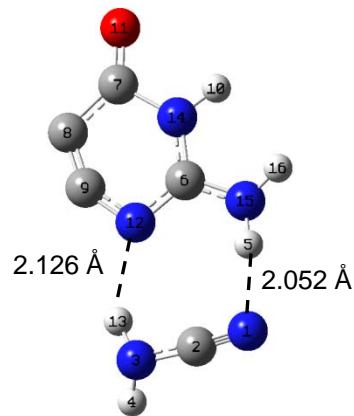
8

$G = -541.126260$

$\mu = 2.99 \text{ D}$

$\nu_1 = 38.73 \text{ cm}^{-1}$

MP2/6-311++G(2d,2p)



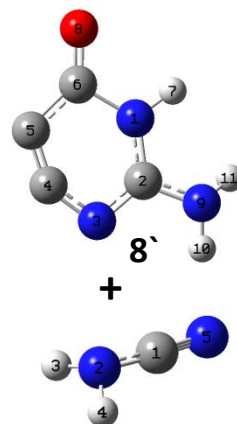
8

$G = -541.126260$

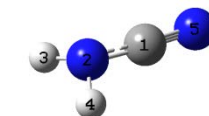
$\mu = 2.99 \text{ D}$

$\nu_1 = 38.73 \text{ cm}^{-1}$

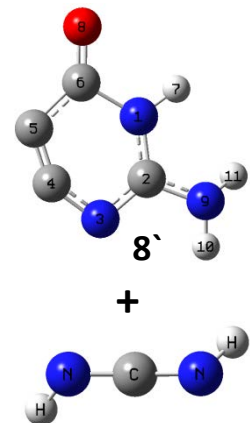
10.78



+



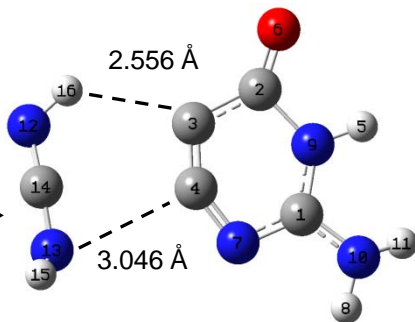
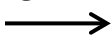
$G = -541.109074$



+

$G = -541.102331$

-3.77



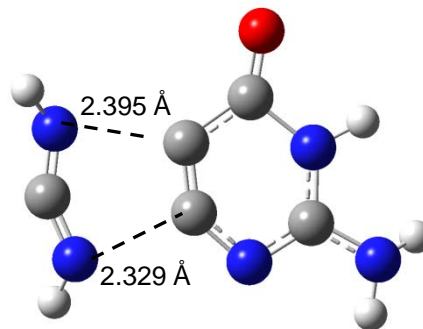
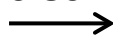
9

$G = -541.108340$

$\mu = 6.39 \text{ D}$

$\nu_1 = 16.96 \text{ cm}^{-1}$

6.80



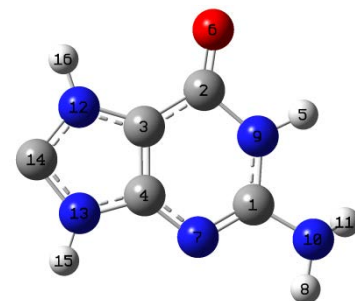
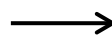
TS_{9,10}

$G = -541.102291$

$\mu = 3.14 \text{ D}$

$\nu_1 = -182.06 \text{ cm}^{-1}$

-77.42



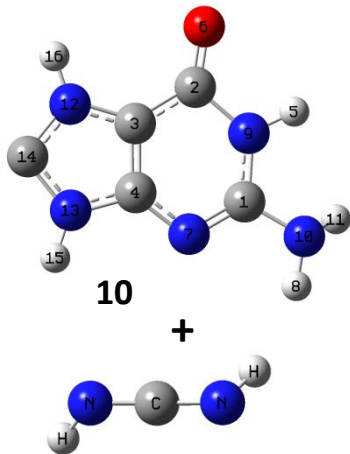
10

$G = -541.225662$

$\mu = 5.55 \text{ D}$

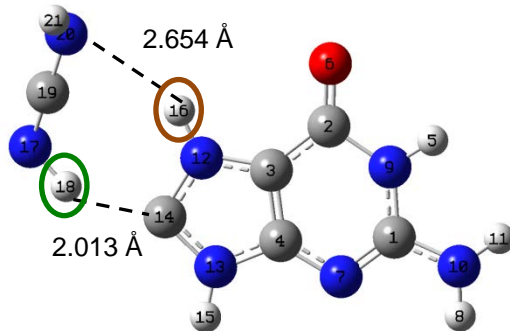
$\nu_1 = 143.27 \text{ cm}^{-1}$

MP2/6-311++G(2d,2p)



G = -689.651040

-8.34

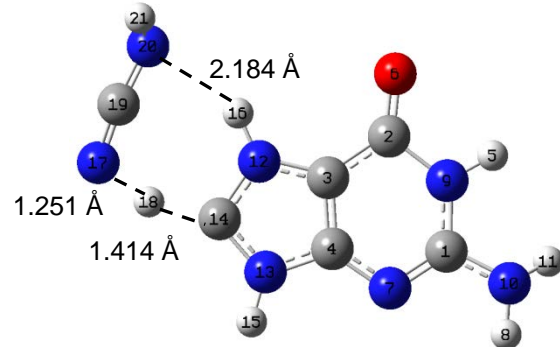


G = -689.664337

$\mu = 7.82$ D

$\nu_1 = 27.35$ cm⁻¹

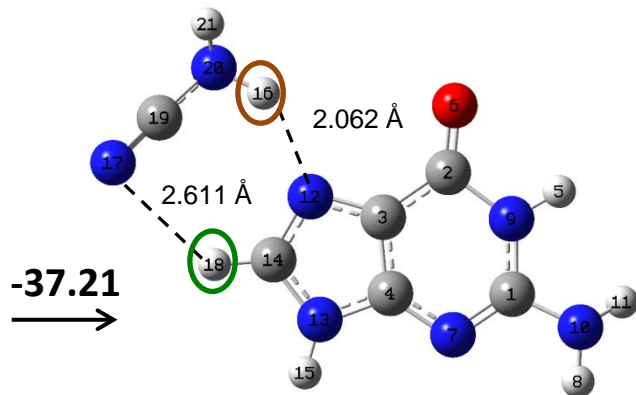
1.89



G = -689.661326

$\mu = 10.65$ D

$\nu_1 = -1112.70$ cm⁻¹



-37.21

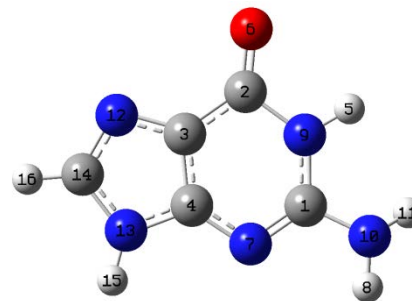
12

G = -689.720623

$\mu = 7.36$ D

$\nu_1 = 22.50$ cm⁻¹

11.32

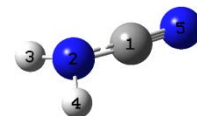


G = -541.270455

$\mu = 6.28$ D

$\nu_1 = 137.12$ cm⁻¹

+

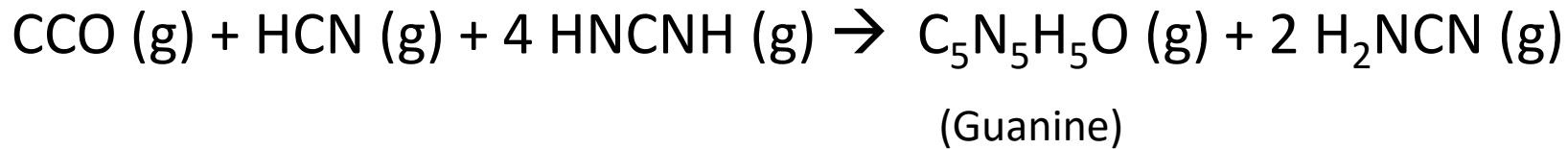


G = -148.432121

$\mu = 4.41$ D

$\nu_1 = 397.84$ cm⁻¹
19

MP2/6-311++G(2d,2p)



$$\Delta_R G^\circ(\text{guanine}) = -228.2 \text{ kcal mol}^{-1}$$

$$\Delta_R G^\circ(\text{adenine}) = -142.2 \text{ kcal mol}^{-1}$$



$$E_a(\text{HNCNH}) = 78.7 \text{ kcal mol}^{-1}$$

Conclusion

- 1 – A new mechanism for the formation of guanine in gas phase was proposed considering the species observed in the ISM: CCO, HCN and HNCNH.
- 2 – The new mechanism has **6 steps** and involves only bimolecular reactions.
- 3 – HNCNH plays a key role for H-bond assisted proton transfer process.
- 4 – We propose first the formation of the 6-membered heterocyclic ring and then the formation of the 5-membered ring.
- 5 – The rate determinant step involves a energy barrier *c.a.* 36 kcal mol⁻¹.
- 6 – This new mechanism is very exergonic, $\Delta G_R = -228.24$ kcal mol⁻¹ .
- 7 – Formation of guanine may be more spontaneous than adenine at the ISM .

Acknowledgments

