

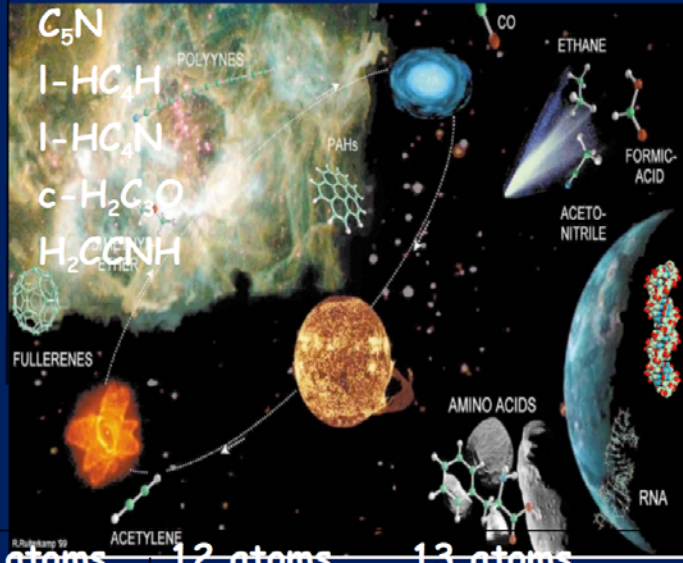
INTERSTELLAR MOLECULES : The Value of Quantum Chemistry

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Theoretical and Computational Chemistry Center.*



2 atoms		3 atoms		4 atoms	5 atoms	6 atoms	7 atoms	8 atoms
H ₂	NaCl	C ₃	N ₂ H ⁺	c-C ₃ H	C ₅	C ₅ H	C ₆ H	CH ₃ C ₃ N
AlF	OH	C ₂ H	N ₂ O	l-C ₃ H	C ₄ H	l-H ₂ C ₄	CH ₂ CHCN	HCOOCH ₃
AlCl	PN	C ₂ O	NaCN	C ₃ N	C ₄ Si	C ₂ H ₄	CH ₃ C ₂ H	CH ₃ COOH
C ₂	SO	C ₂ S	SO ₂	C ₃ O	l-C ₃ H ₂	CH ₃ CN	HC ₅ N	C ₇ H
CH	SO ⁺	CH ₂	c-SiC ₂	C ₃ S	c-C ₃ H ₂	CH ₃ NC	CH ₃ CHO	H ₂ C ₆
CH ⁺	SiO	HCN	CO ₂	C ₂ H ₂	CH ₂ CN	CH ₃ OH	CH ₃ NH ₂	CH ₂ OHCHO
CN	SiS	HCO	NH ₂	HCCN	CH ₄	CH ₃ SH	c-C ₂ H ₄ O	l-HC ₆ H
CO	CS	HCO ⁺	CO ₂ ⁺	HCNH ⁺	HC ₃ N	HC ₃ NH ⁺	H ₂ CCHOH	CH ₂ CHCHO
CO ⁺	HF	HCS ⁺	H ₃ ⁺	HNCO	HC ₂ NH	HC ₂ CHO	C ₆ H ⁻	CH ₂ CCHCN
CP	SH	HOC ⁺	H ₂ D ⁺ , HD ₂ ⁺	HNCS	HCOOH	NH ₂ CHO		NH ₂ CH ₂ CN
SiC	HD	H ₂ O	SiCN	HOCO ⁺	H ₂ CNH			
HCl	FeO?	H ₂ S	AlNC	H ₂ CO	H ₂ C ₂ O			
KCl	O ₂ ?	HNC	OCS	H ₂ CN	H ₂ NCN			
NH	CF ⁺	HNO	HCP	H ₂ CS	HNC ₃			
NO	SiH?	MgCN	CCP	H ₃ O ⁺	SiH ₄			
NS	PO	MgNC		NH ₃	H ₂ COH ⁺			
				c-SiC ₃	C ₄ H ⁻			
				CH ₃	CNCHO			



9 atoms		10 atoms		11 atoms	12 atoms	13 atoms
CH ₃ C ₄ H	CH ₃ CH ₂ OH	SiNC	CH ₃ C ₅ N	HC ₉ N	C ₆ H ₆	HC ₁₁ N
CH ₃ CH ₂ CN	HC ₇ N	CH ₃ C(O)NH ₂	(CH ₃) ₂ CO	CH ₃ C ₆ H	C ₂ H ₅ OCH ₃	
(CH ₃) ₂ O	C ₈ H ⁻	CH ₂ CHCH ₃				

What's it Good For?

- Computational chemistry is a rapidly growing field in chemistry.
 - Computers are getting faster.
 - Algorithms and programs are maturing.
- Some of the almost limitless properties that can be calculated with computational chemistry are:
 - Equilibrium and transition-state structures
 - dipole and quadrupole moments and polarizabilities
 - Vibrational frequencies, IR and Raman Spectra
 - NMR spectra
 - Electronic excitations and UV spectra
 - Reaction rates and cross sections
 - thermochemical data

∞ Basis Set

Complete Basis Set Limit

Exact Solution

Basis Set Expansion

HF
Limit

*Interaction between basis set and correlation method require proper treatment of **both** for accurate calculations.*

Need to specify method and basis set when describing a calculation

QZ

Typical Calculations

TZ

Full CI

DZ

HF

MP2

CCSD(T)

All possible configurations

Wave Function Expansion

Why not use best available correlation method with the largest available basis set?

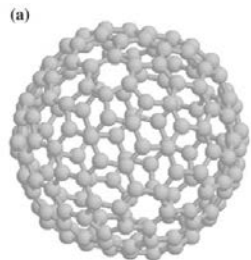
Method	Scaling of Cost
HF	$M^2 - M^3$
MP2	M^5
CCSD	M^6
CCSD(T)	M^7

- A MP2 calculation would be 100x more expensive than HF calculation with same basis set.
- A CCSD(T) calculation would be 10^4 x more expensive than HF calculation with same basis set.
- Tripling basis set size would increase MP2 calculation 243x (3^5).
- Increasing the molecule size 2x (say ethane→butane) would increase a CCSD(T) calculation 128x (2^7).

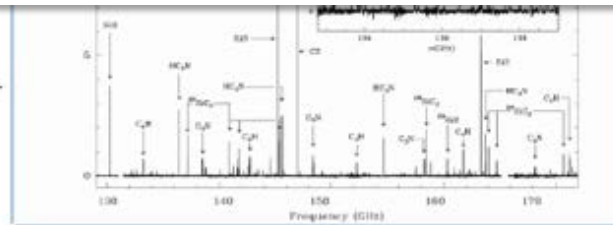
Overview

- Different choices of methods and basis sets can yield a large variation in results.
- It is important to know the errors associated with and limitations of different computational approaches.
- This is important when doing your own calculations, and when evaluating the calculations of others.
- Don't just accept the numbers the computer spits out at face value!

Main Results



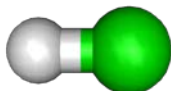
CF^+



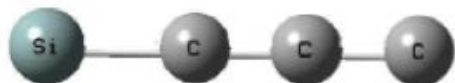
^{28}SiO ^{29}SiO ^{30}SiO



$^{28}SiF^+$ $^{29}SiF^+$ $^{30}SiF^+$



$H^{35}Cl$ $D^{35}Cl$ $T^{35}Cl$ $H^{37}Cl$ $D^{37}Cl$ $T^{37}Cl$



1-SiCCC 1-CSiCC

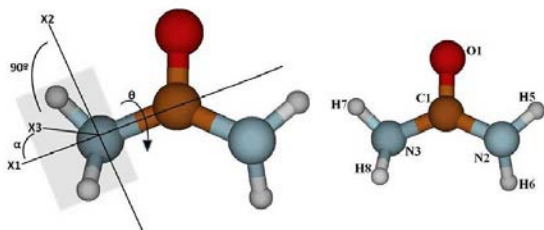
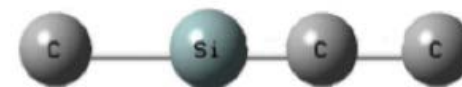
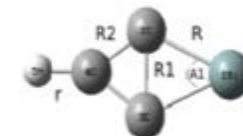
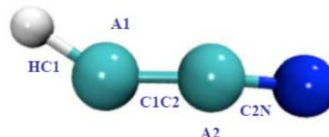


Figure 1. Structural parameters of Urea, definition of NH_2 -torsion (θ) and NH_2 -wagging (α) coordinates.



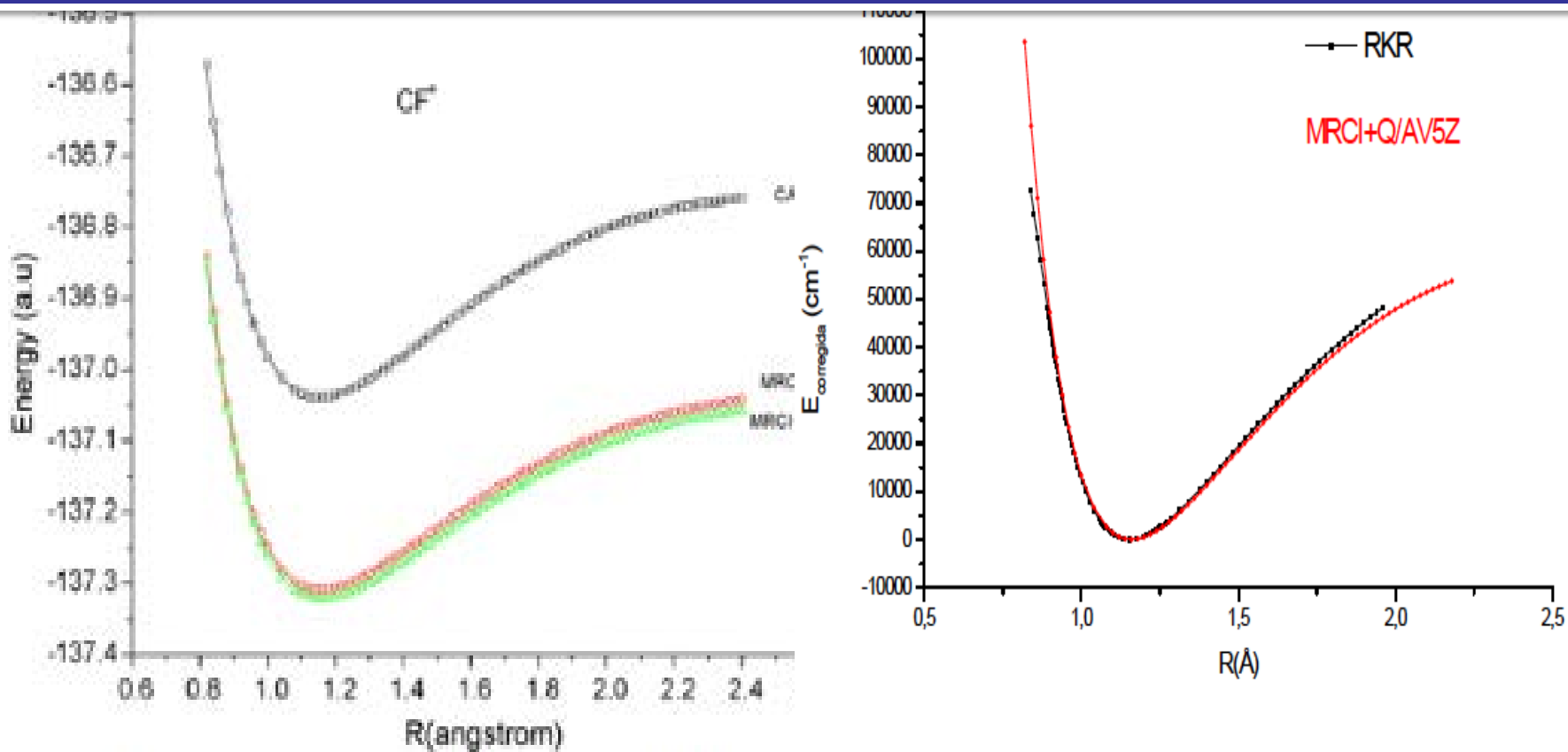


Fig. 1. Potential energy functions of CF^+ , using aug-cc-pV5Z basis set.

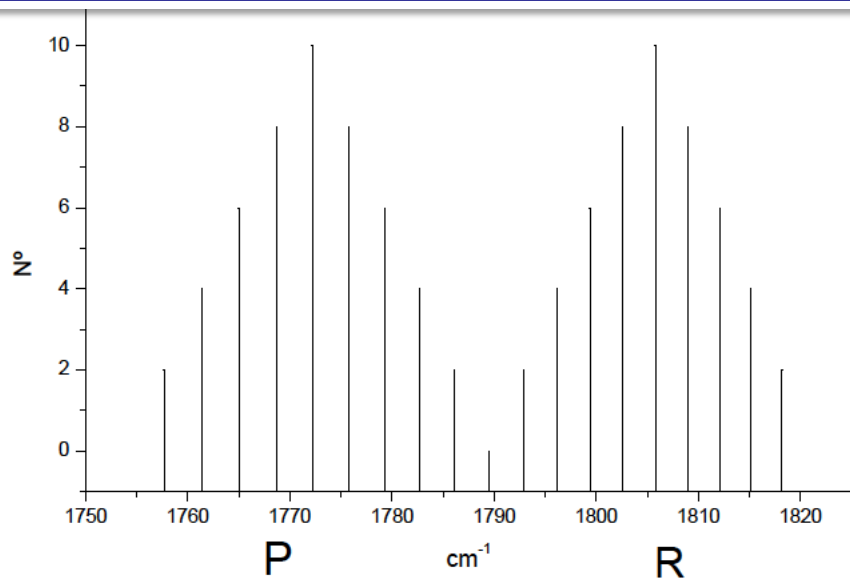
$\Delta ECASSCF-MRCI = 7.33$ eV (169.203 kcal/mol),

$\Delta EMRCI-MRCI+Q = 0.379$ eV

(8.748 kcal/mol)

MRCI+Q/aug-cc-pV5Z para el ión CF^+ versus la curva calculada con el método Rydberg-Klein-Rees (RKR)

Results



$B_e = 1.71662 \text{ cm}^{-1} \text{ (MP)}$.

Dunham Coeff.

$Y_{01} = 1.71013 \text{ cm}^{-1}$

Be (Gruebele)

$Y_{01} = 1.72041 \text{ cm}^{-1}$.

J1/J	0	1	2	3	4	5	6	7	8	9
0	1789.5	1792.9(R0)								
1	1786.1(P1)		1796.2(R1)							
2		1782.7(P2)		1799.4(R2)						
3			1779.3(P3)		1802.6(R3)					
4				1775.8(P4)		1805.8(R4)				
5					1772.2(P5)		1809.0(R5)			
6						1768.7(P6)		1812.1(R6)		
7							1765.0(P7)		1815.1(R7)	
8								1761.4(P8)		1818.2(R8)
9										1757.7(P9)

v	B_v
0	1.69066
1	1.67017
2	1.65137
3	1.63629
4	1.61299
5	1.53989

Rotational transitions

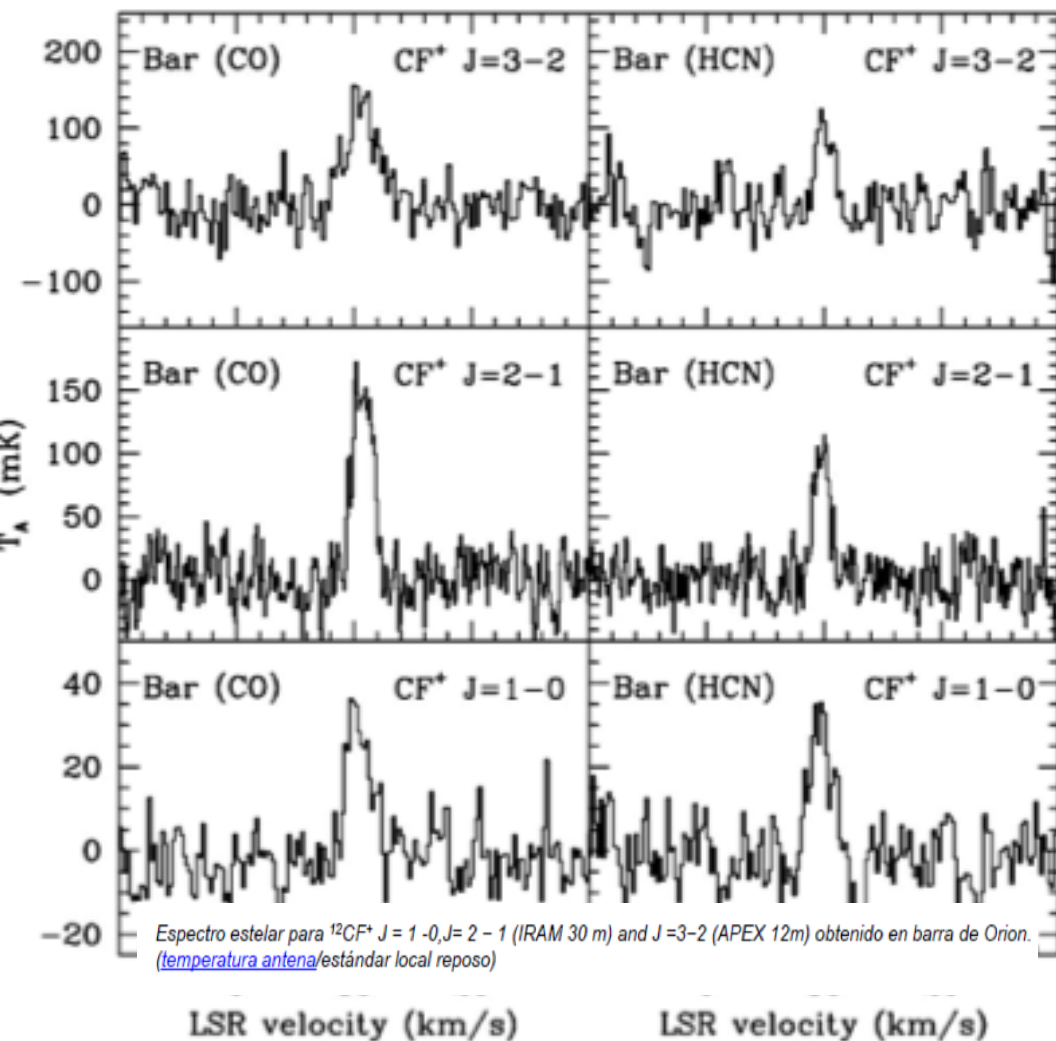
Our results

Obs.

304.2 [GHz] 307.7 [GHz]

203.1 [GHz] 205.2 [GHz]

102.0 [GHz] 102.6 [GHz]

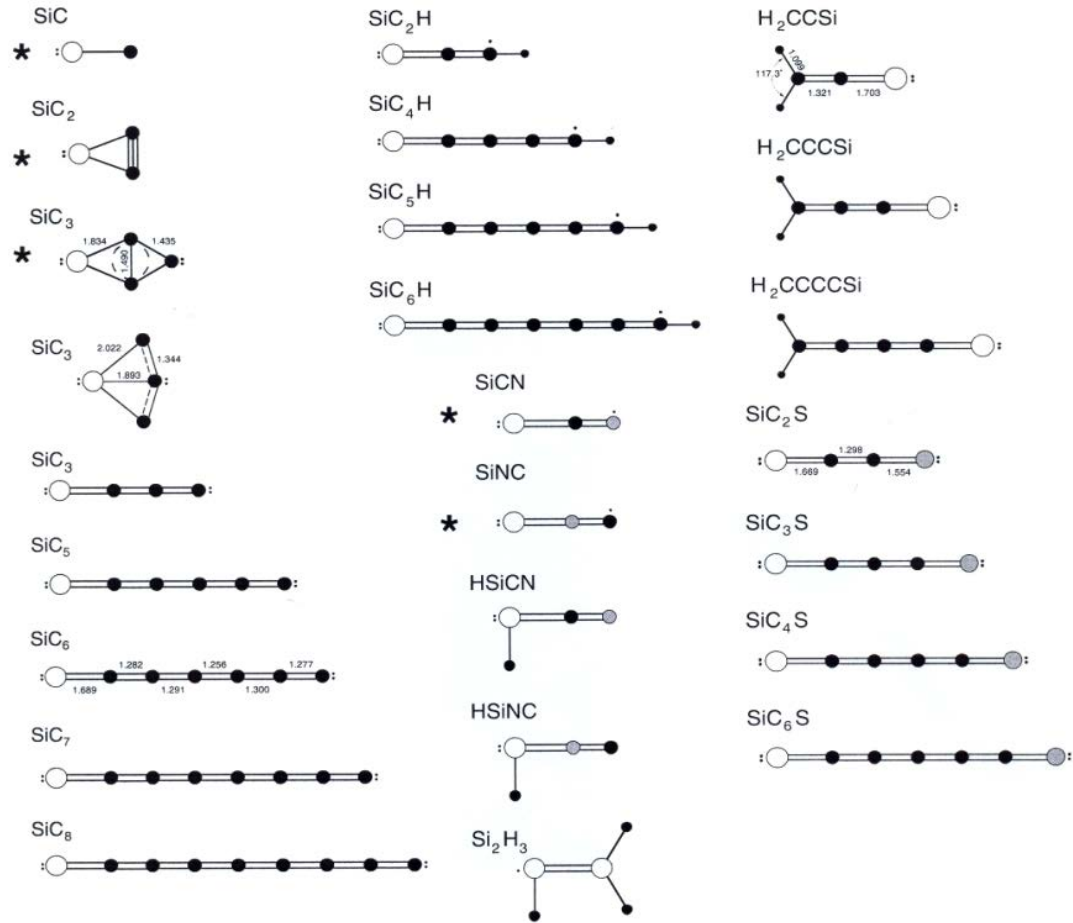
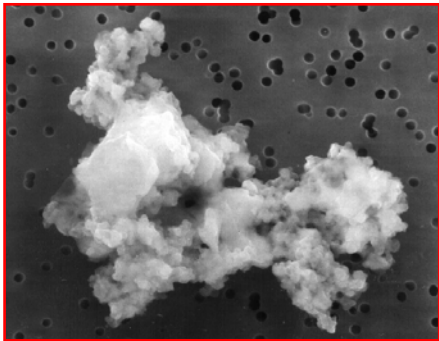


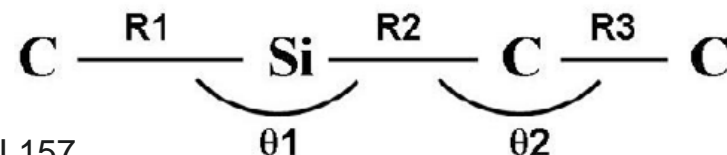
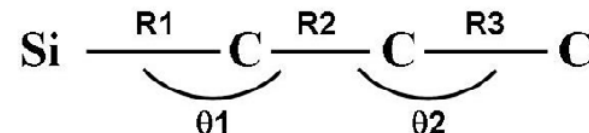
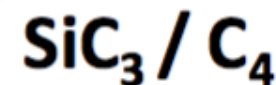
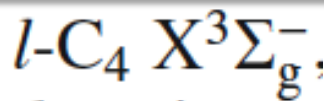
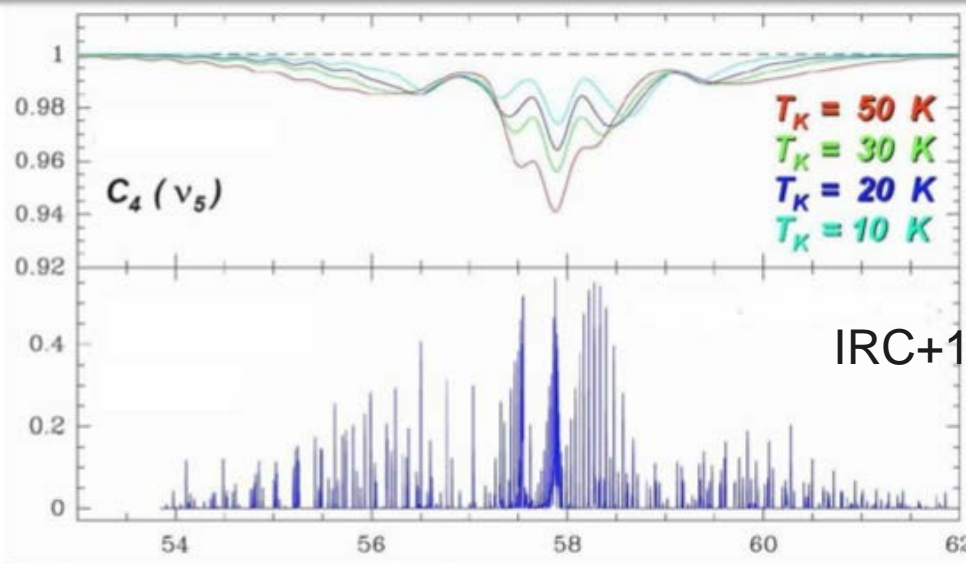
$$\nu = 0, J_2 - J_1 = 1$$

N.Inostroza, J.R.Letelier, P.Fuentealba, M.L.Senent,
Spectrochimica Acta Part A-Molec. and Biomolecular Spectroscopy, 71, 798 (2008).

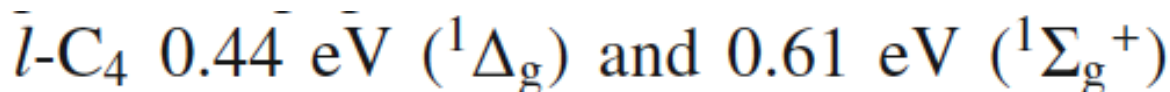
Silicon Carbon Molecules

Molecular Astrophysics:
Silicon carbon molecules
 have been identified in gas
 phase. (Si is a major
 constituent of interstellar
 dust)





Cernicharo, J., Goicochea, J. R., & Benilan, Y. 2002, ApJ, 580, L157



State	CASSCF/cc-pVTZ ^a	MRCI+Q/cc-pVTZ ^b	MCQDPT ^c	Electron configuration
$X^3\Sigma^-$	0.0	0.0	0.0	$(11\sigma)^2(2\pi)^4(3\pi)^2$
$^1\Delta$	0.30	0.29	0.40	$(11\sigma)^2(2\pi)^4(3\pi)^2$
$^1\Sigma^+$	0.44	0.46	0.46	$(11\sigma)^2(2\pi)^4(3\pi)^2$

N. Inostroza, M.L. Senent, M. Hochlaf, Astronomy & Astrophysics 486, 1047 (2008)

		RCCSD(T) <i>cc-pVDZ</i> ^a	RCCSD(T) <i>cc-pVTZ</i> ^b	UCCSD(T) <i>cc-pVTZ</i> ^b	CASSCF <i>cc-pVDZ</i> ^a	CASSCF <i>cc-pVTZ</i> ^b	CASSCF <i>cc-pVQZ</i> ^b	MCSCF 6-31G(d) ^c	CCSD(T) <i>cc-pVQZ</i> ^d	MP2 6-31G(d) ^e	CISD DZP ^f	<i>mω</i> ^g
R_1	(Si-CCC str)	1.7662	1.7447	1.7427	1.7531	1.7410	1.7197	1.74	1.7249	1.732	1.722	
R_2	(SiC-CC str)	1.3299	1.2959	1.2954	1.2966	1.2841	1.2956	1.29	1.2899	1.297	1.298	
R_3	(SiCC-C str)	1.3120	1.3108	1.3133	1.3111	1.3003	1.2987	1.31	1.3062	1.3111	1.3070	
μ						4.0293	4.4064					
B_e		2636.4	2705.49	2706.86	2691.68	2736.88	2753.16					2747.7085
B_0		2637.8			2690.45							
$D_e \times 10^{-6}$		221.529			218.844							255.23
$\omega_1(\sigma)$	(C-C-C stretching)	2016.2	2040	1971	2070.0	2060	2005 ($I = 152$)	2055		2003		
$\omega_2(\sigma)$	(Si-C stretching)	1368.8	1391	1318	1357.1	1365	1366 ($I = 11$)	1342		1332		
$\omega_3(\sigma)$	(C-C-C stretching)	593.8	612	606	611.3	621	639 ($I = 4$)	613		629		
$\omega_4(\pi)$	(trans-bending)	379.8	384	381	443.1	455	463 ($I = 9$)	382		380		
$\omega_5(\pi)$	(cis-bending)	147.2	147	147	166.9	175	166 ($I = 3$)	150		151		

^a This work. Values derived from our 6D PES; ^b this work. Values obtained using standard approaches implemented in MOLPRO. Refs. (Eckert et al. 1997) (Rauhut et al. 1999); ^c Ref. (Rintelman et al. 2001); ^d Ref. (Sattelmeyer et al. 2002); ^e Ref. (Gomei et al. 1997); ^f Ref. (Alberts et al. 1990); ^g Ref. (McCarthy et al. 2000).

N.Inostroza, M.L. Senent, M. Hochlaf, Astronomy & Astrophysics 486, 1047 (2008)

CHARACTERIZATION of the anion SiC_3H^-

The hydrogen-bearing silicon carbide radicals SiC_nH are
isovalent to C_{n+1}H species.

SiC_3H^- isovalent to C_4H^-

C_4H^-

has been one of the first anions detected.

Cernicharo, J. et.al, ApJ. 2002

Neutral C_4H was detected 20 years early.

<i>neutral</i>		μ	<i>anion</i>		μ	E_a
CSi	($X^3\Pi$)	0.8831	l-CSi⁻	($X^2\Sigma^+$)	0.6885	2.42
c-C₂Si	(X^1A_1)	2.9587	l-C₂Si⁻	($X^2\Pi$)	3.8003	1.41
c-C₃Si	(X^1A_1)	3.8671	l-C₃Si⁻	($X^2\Pi$)	4.1481	2.49
l-C₃Si	($X^3\Sigma^-$)	4.4016			--	2.89
l-C₄Si	($X^1\Sigma^+$)	6.2111	l-C₄Si⁻	($X^2\Pi$)	5.9792	2.31
l-C₅Si	($X^3\Sigma^-$)	6.4927	l-C₅Si⁻	($X^2\Pi$)	6.0682	3.30
l-SiCH	($X^2\Pi$)	0.5771	l-SiCH⁻	($X^1\Sigma^+$)	0.3089	3.88
l-SiC₂H	($X^2\Pi$)	1.1201	l-SiC₂H⁻	($X^3\Sigma^-$)	4.2681	1.32
l-SiC₃H	($X^2\Pi$)	1.1074	l-SiC₃H⁻	($X^1\Sigma^+$)	2.9949	2.70
l-SiC₄H	($X^2\Pi$)	1.3061	l-SiC₄H⁻	($X^3\Sigma^-$)	7.3735	1.69
l-SiC₅H	($X^2\Pi$)	0.5122	l-SiC₅H⁻	($X^1\Sigma^+$)	4.8991	2.98

<i>neutral</i>		μ	<i>anion</i>		μ	E_a
CSi	(X ³ Π)	0.8831	l-CSi⁻	(X ² Σ ⁺)	0.6885	2.42
c-C₂Si	(X ¹ A ₁)	2.9587	l-C₂Si⁻	(X ² Π)	3.8003	1.41
c-C₃Si	(X ¹ A ₁)	3.8671	l-C₃Si⁻	(X ² Π)	4.1481	2.49
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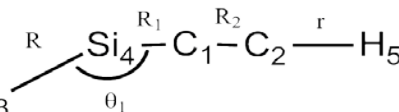
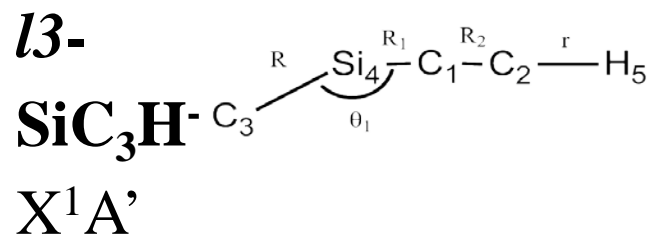
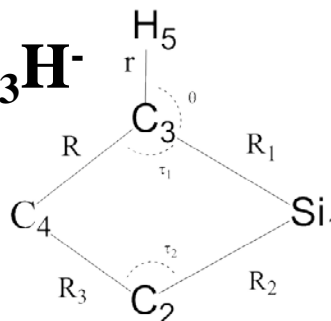
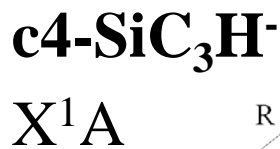
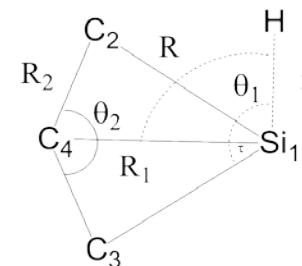
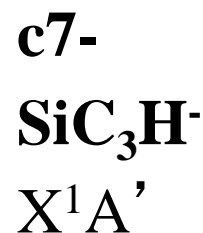
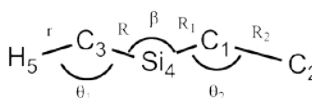
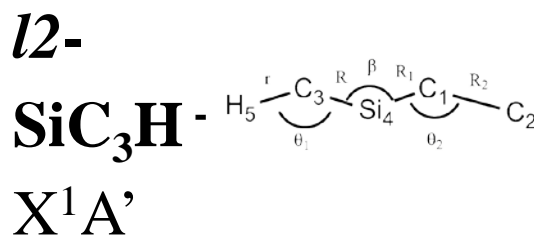
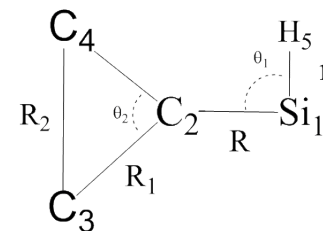
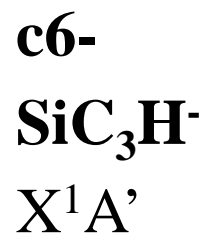
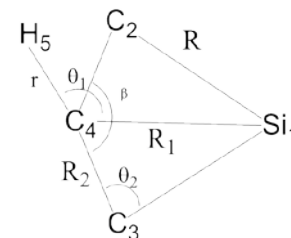
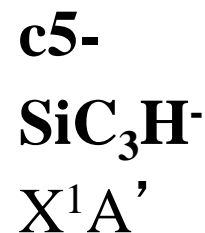
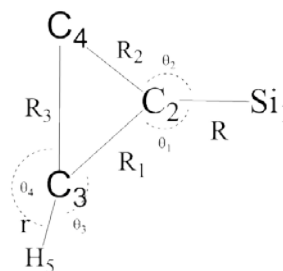
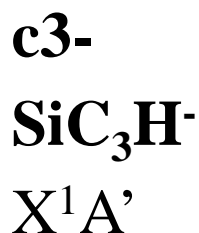
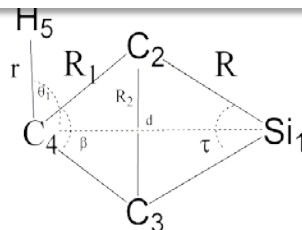
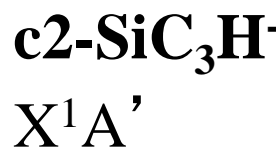
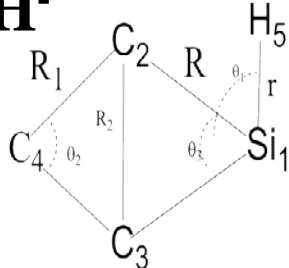
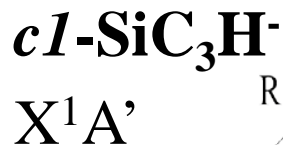
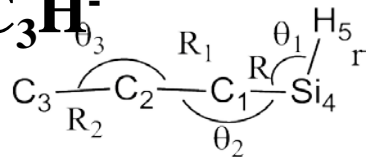
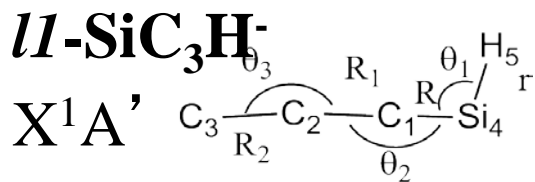
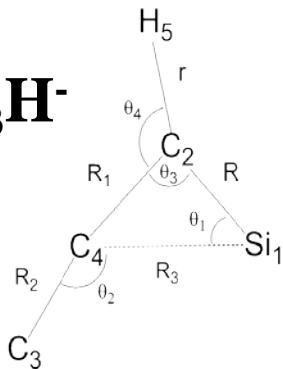
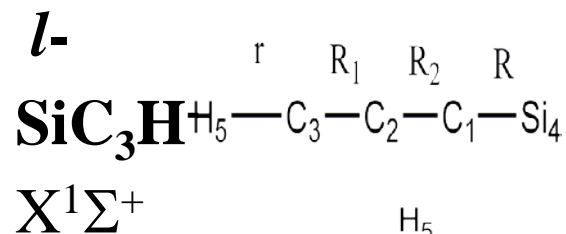
<i>neutral</i>		μ	<i>anion</i>		μ	E_a
CSi	(X ³ Π)	0.8831	l-CSi⁻	(X ² Σ ⁺)	0.6885	2.42
c-C₂Si	(X ¹ A ₁)	2.9587	l-C₂Si⁻	(X ² Π)	3.8003	1.41
c-C₃Si	(X ¹ A ₁)	3.8671	l-C₃Si⁻	(X ² Π)	4.1481	2.49
l-C₃Si	(X ³ Σ ⁻)	4.4016			--	2.89
l-C₄Si	(X ¹ Σ ⁺)	6.2111	l-C₄Si⁻	(X ² Π)	5.9792	2.31
l-C₅Si	(X ³ Σ ⁻)	6.4927	l-C₅Si⁻	(X ² Π)	6.0682	3.30
l-SiCH	(X ² Π)	0.5771	l-SiCH⁻	(X ¹ Σ ⁺)	0.3089	3.88
l-SiC₂H	(X ² Π)	1.1201	l-SiC₂H⁻	(X ³ Σ ⁻)	4.2681	1.32
l-SiC₃H	(X ² Π)	1.1074	l-SiC₃H⁻	(X ¹ Σ ⁺)	2.9949	2.70
l-SiC₄H	(X ² Π)	1.3061	l-SiC₄H⁻	(X ³ Σ ⁻)	7.3735	1.69
l-SiC₅H	(X ² Π)	0.5122	l-SiC₅H⁻	(X ¹ Σ ⁺)	4.8991	2.98

$E_a \rightarrow$ RCCSD(T)-F12A/aug-cc-pVTZ

<i>neutral</i>		μ	<i>anion</i>		μ	E_a	
CSi	(X ³ Π)	0.8831	l-CSi ⁻	(X ² Σ ⁺)	11.5921	2.42	
c-C ₂ Si	(X ¹ A ₁)	2.9587	l-C ₂ Si ⁻	(X ² Π)	5.6931	1.41	
c-C ₃ Si	(X ¹ A ₁)	3.8671	l-C ₃ Si ⁻	(X ² Π)	2.4308	2.49	
l-C ₃ Si	(X ³ Σ ⁻)	4.4016				2.89	
l-C ₄ Si	(X ¹ Σ ⁺)	6.2111	l-C ₄ Si ⁻	(X ² Π)	2.2349	2.31	
l-C ₅ Si	(X ³ Σ ⁻)	6.4927	l-C ₅ Si ⁻	(X ² Π)	5.2765	3.30	
l-SiCH	(X ² Π)	0.577 1	l-SiCH⁻	→	0.3089	3.88	C₂H⁻
l-SiC ₂ H	(X ² Π)	1.120 1	l-SiC ₂ H ⁻	(X ³ Σ ⁻)	1.3651	1.32	
l-SiC ₃ H	(X ² Π)	1.107 4	l-SiC₃H⁻	→	2.9949	2.70	C₄H⁻
l-SiC ₄ H	(X ² Π)	1.306 1	l-SiC ₄ H ⁻	(X ³ Σ ⁻)	10.8913	1.69	
l-SiC ₅ H	(X ² Π)	0.512 2	l-SiC₅H⁻	→	4.8991	2.98	C₆H⁻

<i>neutral</i>		μ	<i>anion</i>		μ	E_a
CSi	(X ³ Π)	0.8831	l-CSi ⁻	(X ² Σ ⁺)	11.5921	2.42
c-C ₂ Si	(X ¹ A ₁)	2.9587	l-C ₂ Si ⁻	(X ² Π)	5.6931	1.41
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l-C ₃ Si	(X ³ Σ ⁻)	4.4016				2.89
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l-SiCH	(X ² Π)	0.577 1	l-SiCH ⁻	(X ¹ Σ ⁺)	4.5084	3.88
l-SiC ₂ H	(X ² Π)	1.120 1	l-SiC ₂ H ⁻	(X ³ Σ ⁻)	1.3651	1.32
l-SiC ₃ H	(X ² Π)	1.107 4	l-SiC₃H⁻	(X ¹ Σ ⁺)	2.9949	2.70
l-SiC ₄ H	(X ² Π)	1.306 1	l-SiC ₄ H ⁻	(X ³ Σ ⁻)	10.8913	1.69
l-SiC ₅ H	(X ² Π)	0.512 2	l-SiC ₅ H ⁻	(X ¹ Σ ⁺)	4.8991	2.98

C₄H⁻



$$B_0 = B_e^{\text{CBS}} + \square B_e^{\text{core}} + \square B_{\text{vib}}$$

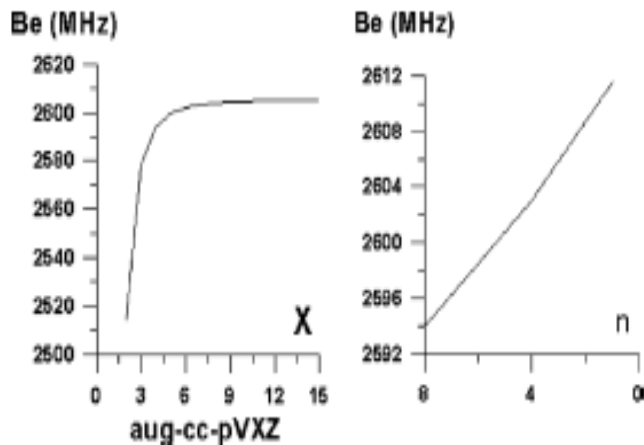
$$\square B_e^{\text{core}} = B_e(\text{aug-cc-pCVQZ}, n=1) - B_e(\text{aug-cc-pVQZ}, n=8)$$

$$\square B_e^{\text{core}} \square 17 \text{ MHz} \quad \square B_{\text{vib}} \square 2 \text{ MHz}$$

$$B_0(l\text{-SiC}_3\text{H}^-) = 2620.74 \text{ MHz} \quad B_0(l\text{-SiC}_3\text{D}^-) = 2459.81 \text{ MHz}$$

$$\square = 2.9707 \text{ Debyes}$$

CASSCF/aug-cc-pV5Z



$$B_e = B_e^{\text{CBS}} + B_e^1 (X+1)^{-3} + B_e^2 (X+1)^{-5} + \dots$$

Basis set	n ^a	<i>l</i> -SiC ₃ H ⁻ B _e	<i>l</i> -SiC ₃ D ⁻ B _e
RCCSD(T)-F12A aug-cc-pVTZ	8	2598.33	2438.24
aug-cc-pVTZ	8	2578.13	2419.55
aug-cc-pVQZ	8	2594.03	2434.29
aug-cc-pV5Z	8	2599.57	2439.36
CBS^b		2605.16	2444.36
aug-cc-pCVQZ	4	2603.00	2442.46
aug-cc-pCVQZ	1	2611.51	2450.45

a) n=number of frozen core orbitals

b) CBS =complete basis set (aug-cc-pV□Z)

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Vertical excitation energies of *l*-SiC₃H⁻

<i>l</i> -SiC ₃ H		<i>l</i> -SiC ₃ H ⁻	
Sym	Er MRCI	Sym	Er MRCI
X²Π	0.0^b	X¹Σ⁺	0.0^d
² Σ ⁺	2.01	¹ Π	3.09
² Δ	5.66	¹ Δ	3.29
² Σ ⁻	5.43	¹ Σ ⁻	3.22
⁴ Σ ⁺	3.43	³ Σ ⁺	2.48
⁴ Δ	3.82	³ Π	3.10
⁴ Σ ⁻	4.19	³ Δ	2.66
⁴ Π	2.21	³ Σ ⁻	2.85

E_a = 2.70 eV

MRCI/aug-cc-pVTZ

c) E_a = -493.205588 a.u.; d) E_a = -403.597248 a.u.

Vertical excitation energies of *l*-SiC₃H⁻

<i>l</i> -SiC ₃ H		<i>l</i> -SiC ₃ H ⁻	
Sym	Er MRCI	Sym	Er MRCI
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² Σ ⁺	2.01	¹ Π	3.09
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⁴ Δ	3.82	³ Π	3.10
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Vertical excitation energies of *l*-SiC₃H⁻

<i>l</i> -SiC ₃ H		<i>l</i> -SiC ₃ H ⁻	
Sym	Er(eV) MRCI	Sym	Er(eV) MRCI
X²Π	0.0^b	X¹Σ⁺	0.0^d
² Σ ⁺	2.01	¹ Π	3.09
² Δ	5.66	¹ Δ	3.29
² Σ ⁻	5.43	¹ Σ ⁻	3.22
⁴ Σ ⁺	3.43	³Σ⁺	2.48
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⁴ Σ ⁻	4.19	³Δ	2.66
⁴ Π	2.21	³Σ⁻	2.85

E_a = 2.70 eV

MRCI/aug-cc-pVTZ

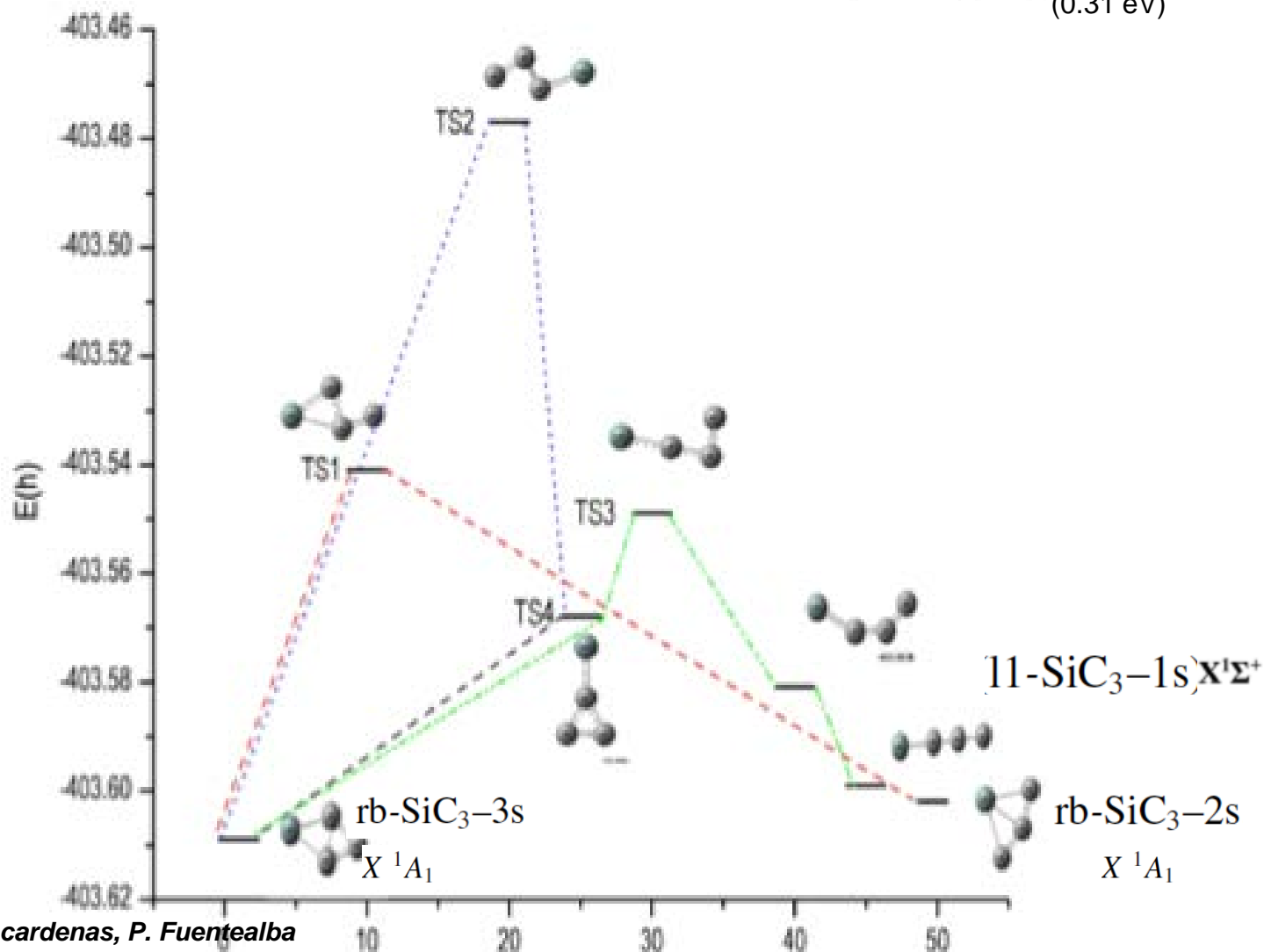
c) E_a = -493.205588 a.u.; d) E_a = -403.597248 a.u.

Relative stability

SiC_3H		E_r B3LYP	E_r CCSD(T)	SiC_3H^+		E_r B3LYP	E_r CCSD(T)
	<i>ll</i> - SiC_3H $C_{\infty v} X^2\Pi$	0.0	0.0		<i>ll</i> - SiC_3H^+ $C_{\infty v} X^1\Sigma^+$	0.55	0.93
	<i>rb1</i> - SiC_3H $C_1 X^2A'$	2.50	2.11		<i>rb1</i> - SiC_3H^+ $C_s X^1A'$	3.58	3.17
	<i>rb2</i> - SiC_3H $C_s X^2A''$	0.62	0.23		<i>rb2</i> - SiC_3H^+ $C_s X^1A'$	0.82	0.86
	<i>rb3</i> - SiC_3H $C_{2v} X^2B_1$	0.69	0.31		<i>rb3</i> - SiC_3H^+ $C_{2v} X^1A_1$	0.00	0.00

(^a) $E(\text{B3LYP/cc-pvTZ}) = -404.303366$ au; (^b) $E(\text{CCSD(T)/cc-pvTZ}) = -403.566239$ a

$X^3\Sigma^- 11\text{-SiC}_3\text{-1t}$ lies $7.2 \text{ kcal mol}^{-1}$
(0.31 eV)



N. Inostroza, C. cardenas, P. Fuentealba

Monthly Notices of the Royal Astronomical Society MNRAS 443, 3127–3133 (2014)

Reaction Channels and spectroscopic constant of astrophysical relevant silicon-bearing molecules SiC₃H⁺ and SiC₃

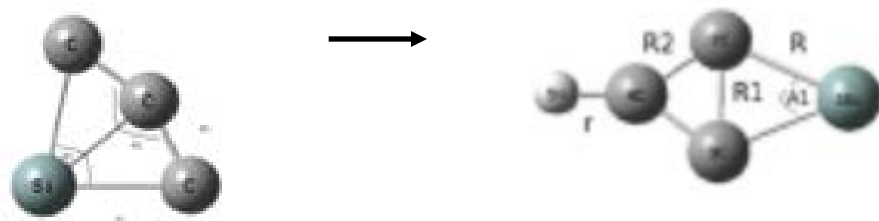
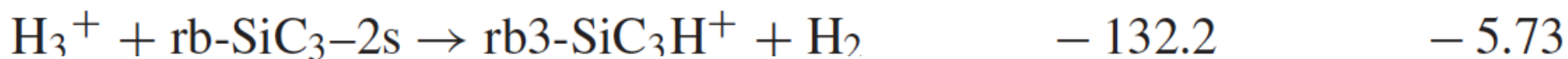
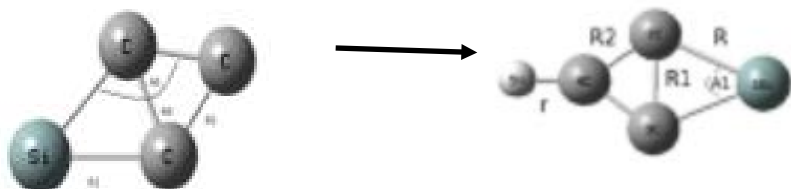
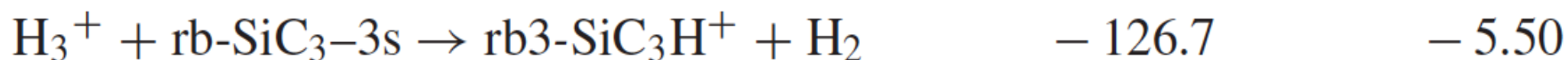
Type Reaction		[ΔE(kcal mol ⁻¹)	ΔE(eV) ^a	[ΔE(kcal mol ⁻¹) ^b	ΔE (eV)] ^b
Charge-exchange (10–300K)	H ⁺ + 1l-SiC ₃ H → 1l-SiC ₃ H ⁺ + H	201.5	8.74	203.8	8.84
	H ⁺ + rb3-SiC ₃ H → rb3-SiC ₃ H ⁺ + H	170.9	7.41	175.4	7.61
	H ⁺ + rb2-SiC ₃ H → rb2-SiC ₃ H ⁺ + H	192.5	8.35	195.2	8.46
Ion-neutral (10–41 000K)	H ₃ ⁺ + rb-SiC ₃ -3s → rb3-SiC ₃ H ⁺ + H ₂	-126.7	-5.50	-129.2	-5.60
	H ₃ ⁺ + rb-SiC ₃ -3s → rb2-SiC ₃ H ⁺ + H ₂	-106.9	-4.63	-109.6	-4.75
	H ₃ ⁺ + rb-SiC ₃ -3s → rb1-SiC ₃ H ⁺ + H ₂	-53.7	-2.33	-46.9	-2.03
	H ₃ ⁺ + rb-SiC ₃ -2s → rb3-SiC ₃ H ⁺ + H ₂	-132.2	-5.73	-135.9	-5.89
	H ₃ ⁺ + rb-SiC ₃ -2s → rb2-SiC ₃ H ⁺ + H ₂	-112.3	-4.87	-114.5	-4.97
	H ₃ ⁺ + rb-SiC ₃ -2s → 1l-SiC ₃ H ⁺ + H ₂	-110.8	-4.80	-115.0	-4.99
	H ₃ ⁺ + 1l-SiC ₃ -1t → rb3-SiC ₃ H ⁺ + H ₂	-132.6	-5.75	-135.4	-5.87
	H ₃ ⁺ + 1l-SiC ₃ -1t → rb2-SiC ₃ H ⁺ + H ₂	-112.7	-4.89	-113.9	-4.94
	H ₃ ⁺ + 1l-SiC ₃ -1t → 1l-SiC ₃ H ⁺ + H ₂	-111.2	-4.82	-114.4	-4.96
Dissociative recombination (10–300K)	rb3-SiC ₃ H ⁺ + e ⁻ → rb-SiC ₃ -3s + H	232.9	10.09	235.1	10.19
	rb2-SiC ₃ H ⁺ + e ⁻ → rb-SiC ₃ -2s + H	218.5	9.47	220.5	9.56
	rb2-SiC ₃ H ⁺ + e ⁻ → rb-SiC ₃ -3s + H	212.9	9.23	215.6	9.35
	1l-SiC ₃ H ⁺ + e ⁻ → 1l-SiC ₃ -1s + H	290.1	12.6	281.3	12.2

^acalculated at CCSD(T)/cc-pvTZ and ^bCCSD(T)/6-311g(d,p) level of theory.

Reaction Channels and spectroscopic constant of astrophysical relevant silicon-bearing molecules SiC₃H⁺ and SiC₃H

Charge-exchange (10–300K)	[$\Delta E(\text{kcal mol}^{-1})$]	$\Delta E(\text{eV})^a$
$\text{H}^+ + 11\text{-SiC}_3\text{H} \rightarrow 11\text{-SiC}_3\text{H}^+ + \text{H}$	201.5	8.74
$\text{H}^+ + \text{rb3-SiC}_3\text{H} \rightarrow \text{rb3-SiC}_3\text{H}^+ + \text{H}$	170.9	7.41
$\text{H}^+ + \text{rb2-SiC}_3\text{H} \rightarrow \text{rb2-SiC}_3\text{H}^+ + \text{H}$	192.5	8.35

Ion-neutral (10–41 000K)

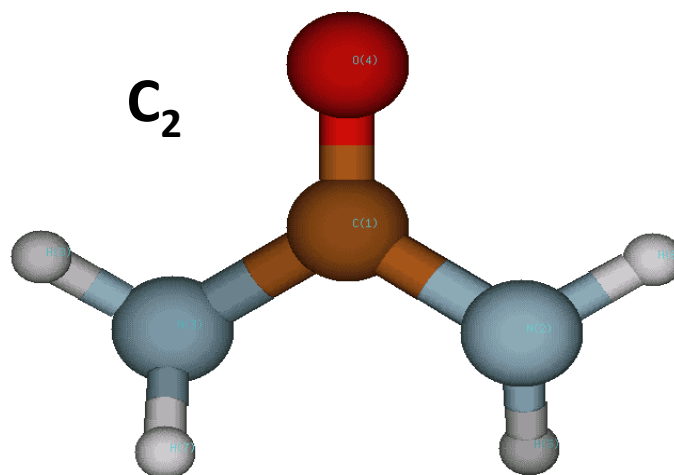


Large amplitude vibrations of Urea in gas phase

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Departamento de Química y Física Teóricas, Instituto de Estructura de la Materia, IEM-CSIC, Serrano 121, Madrid 28006, Spain

Although Urea rarely appears listed as an interstellar species it constitutes an important prebiotic molecule whose astrophysical detection is always expected. A recent and extensive search towards the high mass hot molecular core Sgr B2 (N-LMH) has been performed with the CARMA and IRAM 30 m radio-telescopes observing several line frequencies coincident with Urea transitions [1]. It has been tentatively detected in infrared spectra of interstellar ices [2].



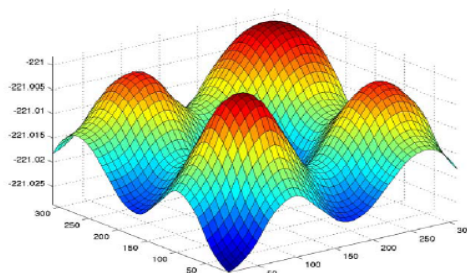
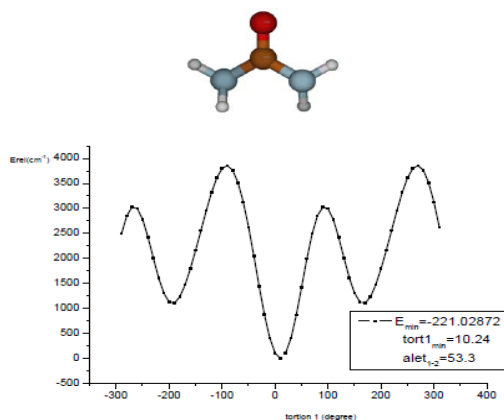
Chemical Physics Letters 524 (2012) 25–31

Potential Energy Surface PES

$$V = \frac{1}{2} \sum_i f_{ij} Q_i Q_j + \frac{1}{6} \sum_i \sum_j \sum_k f_{ijk} Q_i Q_j Q_k + \frac{1}{24} \sum_i \sum_j \sum_k \sum_l f_{ijkl} Q_i Q_j Q_k Q_l + \dots$$

RCCSD(T)-F12/ cc-pVTZ-F12

UREA



GRID

1409 geometries :

Bond distances $R^{\text{ref}} + 0.03 \text{ \AA} \geq R \geq R^{\text{ref}} - 0.03 \text{ \AA}$

Torsional angles $\theta^{\text{ref}} + 5.0^\circ \geq \theta \geq \theta^{\text{ref}} - 5.0^\circ$

Planar bending angles = $\tau + 5.0^\circ$

FIT

$R^2 = 1.0$, $\sigma = 0.4 \text{ cm}^{-1}$

FIT-ESPEC (Senent 2007)

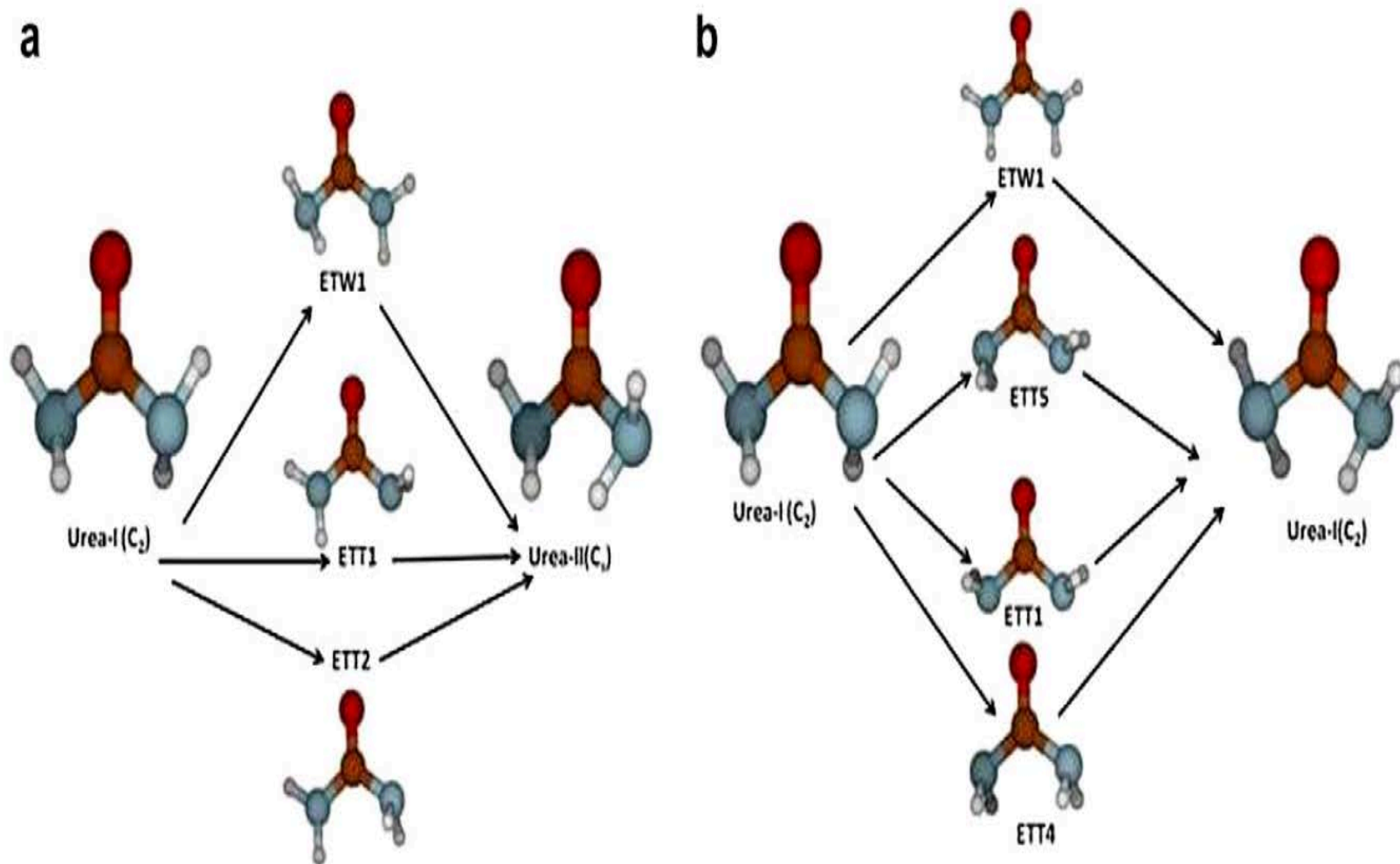
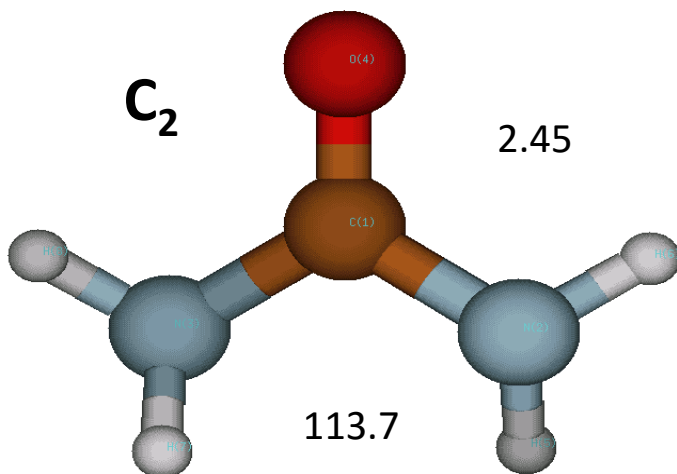


Figure 2. Conversion channels and transition states for the processes (a) Urea-I \rightarrow Urea-II; (b) Urea-I \rightarrow Urea-I.

2794 cm^{-1} (torsional barrier) and 395 cm^{-1} (inversion barrier)

Urea I



Urea II

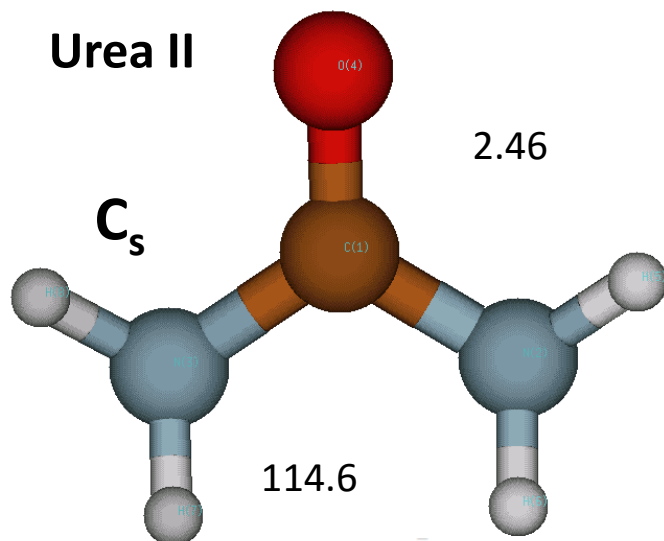
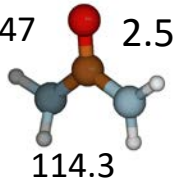
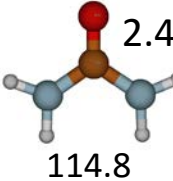
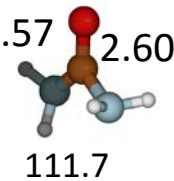
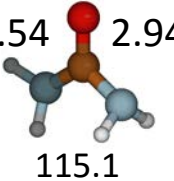
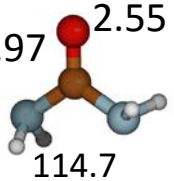
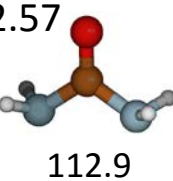
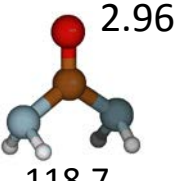


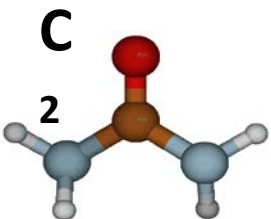
TABLE 1: Total electronic energies (E_a , in a.u.), relative energies (E_R in cm^{-1}), structural parameters (distances in Å; angles in degrees), rotational constants (in MHz) and dipole moments (in Debyes) corresponding to the two conformers of UREA.

	Urea-I (C_{2h})		Urea-II (C_s)	
aug-cc-pVTZ	MP2	CCSD	MP2	CCSD-SYM
E_a	-224.9106098	-224.9217913	-224.9091131	-224.9191171
E_R	0.0	0.0	328.48	510.61
C-O	1.2191	1.2118	1.2206	1.2139
C-N	1.3838	1.3835	1.3789	1.3777
N-C-O	123.16	123.11	122.61	122.85
H5-N-C	117.14	116.8	120.08	117.36
H6-N-C	112.86	112.75	114.11	117.36
θ_1	5.7861	5.8338	176.185	176.82
β_1	38.5597	39.3752	30.3674	29.4000
θ_2	5.7930	5.8349	3.8263	3.1800
β_2	38.5606	39.3762	30.3949	29.4000
A	11136.50	11229.32	11219.06	11244.3954
B	10422.07	10419.66	10393.14	10458.7141
C	5433.06	5456.16	5421.96	5446.2361
μ	4.03	3.95	4.67	4.64

$$\hat{H}_{\text{LAM}} = - \sum_i^n \sum_j^n \left(\frac{\partial}{\partial q_j} \right) B_{ij} \left(\frac{\partial}{\partial q_j} \right) + V(q_1, q_2, \dots, q_n) + V^1(q_1, q_2, \dots, q_n)$$

Urea

							
	ETW1	ETW2	ETT1	ETT2	ETT3	ETT4	ETT5
Sym.	C ₁	C _{2v}	C ₁	C ₁	C _s	C _{2v}	C _{2v}
C-O	1.2208	1.2221	1.2158	1.2109	1.2042	1.2100	1.2002
C-N2	1.3860	1.3709	1.4020	1.4067	1.4442	1.4413	1.4475
C-N3	1.3685	1.3708	1.4020	1.4067	1.4442	1.4413	1.4475
N2-C-O	122.75	122.58	124.16	122.44	122.66	123.53	120.67
N3C-O	122.85	122.58	124.16	122.44	122.66	123.53	120.67
H5-N2-C	112.48	123.56	117.67	117.67	106.60	106.57	108.80
H6-N2-C	117.90	116.97	117.67	117.67	106.60	106.57	108.80
H7-N3-C	123.08	123.55	108.40	111.44	106.60	106.57	108.80
H8-N3-C	117.32	116.98	108.40	111.44	106.60	106.57	108.80
θ ₁	6.4090	0.0	1.247	3.040	270.0	90.0	270.0
β ₁	38.120	0.0	33.862	33.358	62.662	62.601	58.161
θ ₂	1.0780	0.0	90.0	270.0	90.00	90.0	270.0
β ₂	0.000	0.0	55.654	48.895	62.662	62.601	58.161
μ	4.5140	4.7095	2.8386	4.9981	3.2763	0.2148	5.2990
E _a	-224.90921	-224.90834	-224.89205	-224.88606	-224.86408	-224.85917	-224.85917
E _R	306.4	497.7	4072.6	5388.4	10212.2	10660.9	11290.5



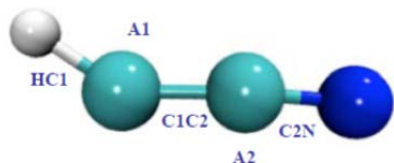
Urea-I (C₂)

Sym	MP2		CCSD(T)	exp	Assign.
	ω	ν	ω	[19]	
A	3723	3560	3685	3559	NH stretch
ν_2 A	3600	3452	3573	3460	NHstretch
ν_3 A	1800	1757	1789	1776	CO stretch
ν_4 A	1635	1592	1639	1604	HNH ben
ν_5 A	1183	1150	1191	1157	HNH ben
ν_6 A	959	942	949	1032	CN stretch
ν_7 A	575	421	602	582.4^b	NH2 wag
ν_8 A	475	472	472		NCN ben
ν_9 A	377	338	377		Torsion
ν_{10} B	3723	3560	3685	3533	NHstretch
ν_{11} B	3595	3450	3570	3434	NHstretch
ν_{12} B	1640	1594	1645	1749	HNH ben
ν_{13} B	1420	1384	1417	1394	CN stretch
ν_{14} B	1057	1011	1064	1157	HNH ben
ν_{15} B	782	726	784	775	CO wag
ν_{16} B	581	556	582	571	NCO ben
ν_{17} B	545	495	547	550.6^b	NH2 wag
ν_{18} B	450	380	450	445.1^b	Torsion
ZPVE	14060	13420	14011		

^a Basis set: aug-cc-pVTZ; ZPVE and LAM frequencies, in bold.

^b Experiments in Ar-matriz [15].

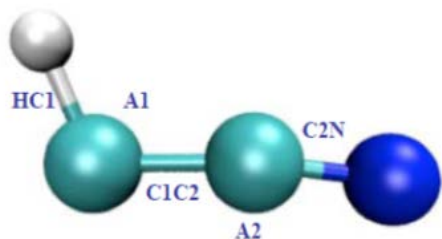
Accurate ab initio quartic force fields of cyclic and bent HC₂N isomers



X³A'' ground state triplet

The grids for each electronic state consisted of 743 distinct geometries and these were used to fit our best QFFs.

CCSD(T) or RCCSD(T) /cc-pVX Z, X = 3,4,5,

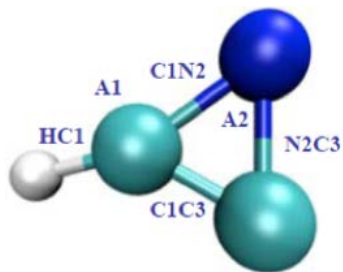


X¹A' bent singlet

$$E(l) = E(TQ5) + E(\text{rel} - \text{nrel}) + E(\text{mtcc} - \text{nmtcc}),$$

scalar relativistic effects

core-correlation correction

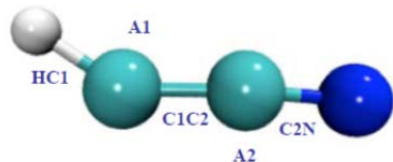


X¹A' cyclic singlet

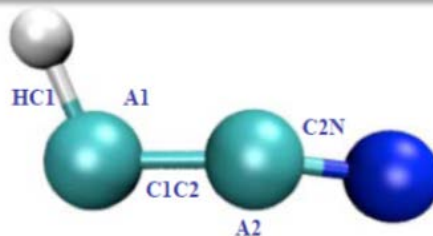
The QFFs were used together with second-order perturbation theory (PT) (SPECTRO) and variational methods (MULTIMODE) to solve the nuclear Schrödinger equation.

natalia inostroza, *Xinchuan Huang*, and *Timothy J. Lee*. *J. Chem. Phys.* 135, 244310 (2011)

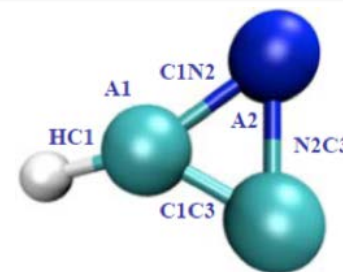
Accurate *ab initio* quartic force fields of cyclic and bent HC₂N isomers



X³A'' ground state triplet



X¹A' bent singlet



X¹A' cyclic singlet

Isomeric energy differences

HC ₂ N	ΔE^a	ΔE^b	μ^c
Ground state triplet	0.0	0.0	3.05
Cyclic singlet	5.7	7.8	3.06
Bent singlet	10.6	11.1	1.71

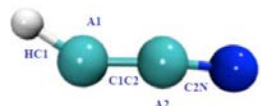
^aEnergies came from the best *ab initio* QFFs, 3-pt(tz,qz,5z)+core+rel.

^bIncludes anharmonic zero-point energies corrections.

^cDipole moments computed at CCSD(T)/cc-pVQZ level of theory.

Relative energies (in kcal/mol)

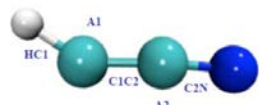
Natalia Inostroza, Xinchuan Huang, and Timothy J. Lee. *J. Chem. Phys.* 135, 244310 (2011)



X³A'' ground state triplet

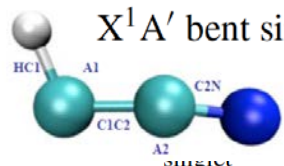
Ground state triplet	PT				VCI			Previous work	
	2-pt (tz,qz)	2-pt (qz,5z)	3-pt (tz,qz,5z)	5z +core+rel	3-pt (tz,qz,5z) +core	3-pt (tz,qz,5z) +core+rel	3-pt (tz,qz,5z) +core+rel	Experiment	Theory
A ₀	2586574	2518363	2496121	2611229	2632677	2614091		4350000 ^b	10938.6 ^e
B ₀	11016	11001	10995	11037	11040	11043		11027 ^b	
C ₀	10951	10935	10928	10972	10975	10979		10986.41 ^a 10986.4 ^b	
D _J	0.0042	0.0042	0.0042	0.0042	0.0042	0.0042		0.0041 ^b	

<u>A₀</u>	2614091	4350000 ^(b)	
<u>B₀</u>	11043	11027 ^(b)	
<u>C₀</u>	10979	10986.41 ^(a) 1098	10938.6 ^(e)
		6.4 ^(b)	



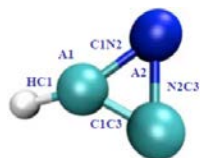
	<u>This Work</u>		<u>Previous Work</u>	
	PT 3-pt (tz,qz,5z) <u>+core+rel</u>	VC 3-pt (tz,qz,5z) <u>+core+rel</u>	Experiment	Theory
$\nu_1(A')$	3243.2	3271.2	3229.0 ^(c) - 3247 ^(d)	3246.66 ^(e) 3245.2 ^(f)
$\nu_2(A')$	1722.8	1615.5 ^{&}	1735 ^(c) -1727, 1735,	1733.71 ^(e) 1851.0 ^(f)
$\nu_3(A')^\#$	1159.3	1177.4	1757 ^(d)	1178.57 ^(e) 1113.9 ^(f)
$\nu_4(A')^{a1}$	276.6	305.8	1178 ^(c)	610.4 ^(f)
$\nu_5(A')$	489.0	556.2	458 ^(c) -383 ^(d) -365 ^(a)	336.2 ^(f)
$\nu_6(A'')$	476.8	561.4	369 ^(c) -187 ^(d) -145 ^(a)	362.1 ^(f)

a) From microwave spectra ref. [23]; b) From microwave spectra ref. [17]; c) From argon matrix IR spectra ref. [16] d) From High resolution infrared spectra ref. [25]; e) MR-ACPF/cc-pVQZ; f) From ref. [34] at CASSCF/DZP



	PT						VCI
	2-pt (tz,qz)	2-pt (qz,5z)	3-pt (tz,qz,5z)	5z +core+rel	3-pt (tz,qz,5z) +core	3-pt (tz,qz,5z) +core+rel	3-pt (tz,qz,5z) +core+rel
A_0	540 126	539 186	538 717	542 306	543 795	543 130	
B_0	11 069	11 052	11 045	11 093	11 098	11 099	
C_0	10 833	10 816	10 809	10 857	10 861	10 863	
HC1	1.1065	1.1067	1.1069	1.1406	1.1048	1.1049	
C1C2	1.3910	1.3920	1.3924	1.3913	1.3881	1.3881	
C2N	1.1765	1.1777	1.1782	1.1755	1.1758	1.1756	
A1(HC1C2)	109.49	109.41	109.38	108.59	109.70	109.64	
A2(C1C2N)	172.44	172.43	172.43	172.46	172.52	172.49	
$10^9 H_J$	-1.9452	-1.9856	-1.9954	-1.9573	-1.9578	-1.9741	
H_K	0.4942	0.4802	0.4749	0.5040	0.5156	0.5083	
$10^6 H_{JK}$	2.4776	2.5099	2.5173	2.4845	2.4918	2.5012	
H_{KJ}	-0.0020	-0.0021	-0.0021	-0.0021	-0.0021	-0.0021	
$10^{11} h_1$	6.7268	6.5112	6.4392	6.7592	6.8086	6.7709	
$10^{10} h_2$	1.5934	1.5959	1.5957	1.5935	1.5925	1.5988	
$10^{11} h_3$	3.2622	3.2423	3.2350	3.2604	3.2545	3.2630	
D_J	0.0050	0.0050	0.0050	0.0051	0.0050	0.0051	
D_K	228.25	225.56	224.45	231.19	233.91	232.41	
D_{JK}	0.9403	0.9373	0.9360	0.9444	0.9461	0.9459	
$10^3 d_1$	-0.0868	-0.0868	-0.0868	-0.0867	-0.0864	-0.0867	
$10^3 d_2$	-0.0214	-0.0213	-0.0212	-0.0214	-0.0213	-0.0213	
$\nu_1(A')^a$	2928.8	2926.7	2925.8	2932.9	2935.3	2933.8	2934.9
$\nu_2(A')^b$	2043.3	2045.1	2045.3	2048.7	2050.9	2050.7	2050.0
$\nu_3(A')$	1040.4	1042.5	1042.8	1044.4	1045.0	1046.9	1046.9
$\nu_4(A')^c$	949.1	956.8	958.3	953.8	957.5	958.4	957.0
$\nu_5(A')$	310.4	321.1	323.3	318.3	325.9	325.1	323.9
$\nu_6(A'')$	443.2	442.6	442.3	445.4	446.1	445.7	444.2

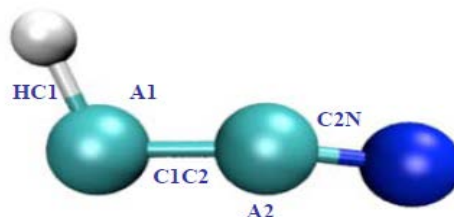
^aFermi resonance type $2\nu_1 = \nu_2 + \nu_4$.


 X^1A' cyclic singlet

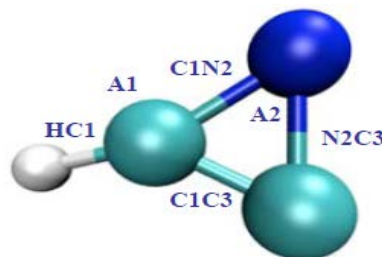
	PT				VCI		
	2-pt (tz,qz)	2-pt (qz,5z)	3-pt (tz,qz,5z)	5z +core+rel	3-pt (tz,qz,5z) +core	3-pt (tz,qz,5z) +core+rel	3-pt (tz,qz,5z) +co re+rel
A_0	40 710	40 599	40 557	40 715	40 755	40 744	
B_0	34 414	34 347	34 322	34 489	34 497	34 505	
C_0	18 584	18 541	18 525	18 607	18 618	18 618	
HC1	1.0798	1.0799	1.0800	1.0787	1.0786	1.0786	
C1N2	1.3005	1.3020	1.3026	1.2998	1.2997	1.2996	
N2C3	1.4152	1.4171	1.4179	1.4148	1.4138	1.4141	
C1C3	1.4066	1.4082	1.4087	1.4052	1.4050	1.4047	
A1(HC1N2)	137.87	137.84	137.84	137.93	137.97	137.94	
A2(C1N2C3)	62.22	62.21	62.20	62.18	62.20	62.18	
$10^7 H_J$	-2.8738	-2.9026	-2.9074	-2.8896	-2.9053	-2.9132	
$10^6 H_K$	18.495	18.533	18.596	18.619	18.153	18.336	
$10^6 H_{JK}$	6.5936	6.6283	6.6488	6.6516	6.5510	6.6001	
$10^6 H_{KJ}$	-21.723	-21.796	-21.876	-21.910	-21.399	-21.605	
$10^{10} h_1$	801.93	792.82	793.75	810.24	781.45	789.14	
$10^{10} h_2$	1886.4	1886.4	1885.4	1891.1	1892.1	1895.6	
$10^{10} h_3$	-347.91	-353.36	-357.60	-359.28	-337.83	-345.77	
D_J	0.0471	0.0469	0.0469	0.0470	0.0472	0.0472	
D_K	0.0904	0.0885	0.0876	0.0882	0.0902	0.0893	
D_{JK}	0.0954	0.0961	0.0966	0.0970	0.0949	0.0960	
d_1	-0.0263	-0.0263	-0.0263	-0.0264	-0.0264	-0.0264	
d_2	-0.0090	-0.0090	-0.0090	-0.0090	-0.0090	-0.0090	
$\nu_1(A')^a$	3122.7	3118.6	3117.4	3126.1	3125.8	3124.7	3126.0
$\nu_2(A')^b$	1571.1	1567.0	1565.7	1572.8	1573.9	1572.5	1576.5
$\nu_3(A')$	1292.0	1288.6	1287.6	1294.4	1294.1	1293.6	1294.1
$\nu_4(A')$	1012.5	1012.0	1011.6	1013.7	1015.9	1015.1	1014.4
$\nu_5(A')$	823.9	821.3	820.4	825.1	825.9	824.9	832.4
$\nu_6(A'')$	899.6	894.5	893.3	898.8	896.4	896.0	901.3

^aFermi resonance type 1 $\nu_1 = 2\nu_2$.

ROVIBRATIONAL SPECTROSCOPIC FOR ISOTOPOLOGUES OF CYCLIC AND BENT SINGLET HC₂N ISOMERS



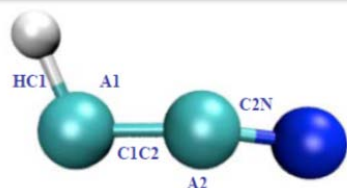
H¹³CCN	HC¹³CN	DCCN	HCC¹⁵N
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DCNC	H¹³CNC	HC¹⁵NC	HCN¹³C
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Natalia Inostroza , R. Fortenberry , X. Huang , and Timothy J. Lee,
The Astrophysical Journal, 778:160 (7pp), 2013 December

ROVIBRATIONAL SPECTROSCOPIC FOR ISOTOPOLOGUES OF CYCLIC AND BENT SINGLET HC₂N ISOMERS



X¹A' bent singlet

CcCR Rotational Constants (MHz), and Fundamental Vibrational Frequencies (cm⁻¹) for the Bent, Singlet HC₂N Isomer Isotopologues

	DCCN		H ¹³ CCN		HC ¹³ CN		HCC ¹⁵ N	
A ₀	307257		537301		542421		543095	
B ₀	10366		10695		11099		10755	
C ₀	10010		10473		10862		10533	
	VPT2 ^a	VCI	VPT2 ^b	VCI	VPT2 ^c	VCI	VPT2 ^b	VCI
ν ₁ (a')	2187.4	2187.2	2925.5	2925.4	2936.1	2932.9	2933.2	2933.5
ν ₂ (a')	2042.3	2041.0	2043.3	2042.4	2016.0 ^d 1948.0 ^d	2015.5 ^d 1950.5 ^d	2031.5	2029.6
ν ₃ (a')	1001.6	1001.3	1035.9	1034.6	1043.1	1042.8	1039.6	1039.3
ν ₄ (a')	797.1	797.7	939.6	938.7	952.8	950.4	956.8	954.1
ν ₅ (a')	311.5	309.7	323.8	320.6	316.8	313.9	323.3	320.2
ν ₆ (a'')	401.7	401.1	444.4	442.4	436.1	434.3	443.0	441.1

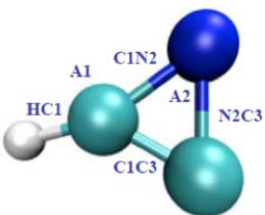
Notes.

^a The ν₂ = 2ν₃, ν₄ = 2ν₆, and ν₆ = 2ν₅ Fermi resonances are included.

^b ν₁ = ν₂ + ν₄, ν₂ = 2ν₃ = 2ν₄ = ν₄ + ν₃, ν₃ = 2ν₆, and ν₄ = 2ν₆ Fermi resonance polyads.

^c Require the ν₁ = ν₂ + ν₄, ν₄ = 2ν₆, and ν₂ = 2ν₃ = 2ν₄ = ν₄ + ν₃ Fermi resonance polyads.

^d These states are coupled at 50%–50% from the ν₂ = ν₄ + ν₃ and ν₂ = 2ν₃ bases.



X¹A' cyclic singlet

CcCR Rotational Constants (MHz) and Fundamental Vibrational Frequencies (cm⁻¹) for the Cyclic, Singlet HC₂N Isomer Isotopologues

	DCCN		H ¹³ CCN		HC ¹³ CN		HCC ¹⁵ N	
A ₀	42505		43549		41784		43783	
B ₀	27137		31172		31934		30843	
C ₀	16508		18106		18038		18034	
	VPT2 ^a	VCI	VPT2 ^b	VCI	VPT2 ^b	VCI	VPT2 ^b	VCI
ν ₁ (a')	2364.2	2364.8	3125.4	3123.3	3146.6	3139.7	3137.9	3135.7
ν ₂ (a')	1530.7	1535.8	1548.0	1551.3	1565.0	1569.4	1552.8	1556.3
ν ₃ (a')	1270.3	1271.0	1272.9	1273.5	1274.4	1274.9	1288.6	1289.2
ν ₄ (a')	961.9	965.6	1011.3	1010.3	1003.1	1002.2	1004.6	1003.6
ν ₅ (a')	665.7	669.6	821.9	829.5	817.5	825.0	820.3	827.8
ν ₆ (a'')	712.3	715.5	888.2	893.4	895.2	900.6	895.8	901.1

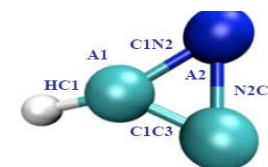
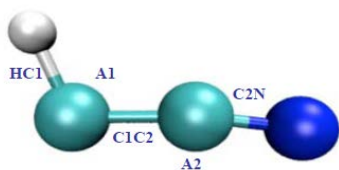
Notes.

^a Fermi resonance ν₁ = 2ν₃, ν₂ = 2ν₆, and ν₃ = 2ν₆.

^b Fermi resonance ν₁ = 2ν₂ and ν₂ = 2ν₅.

Cyanomethylene HCCN

Natalia Inostroza-Pino¹, Partha P. Bera^{2,3}, Xinchuan Huang^{2,4}, and Timothy J. Lee^{2*}

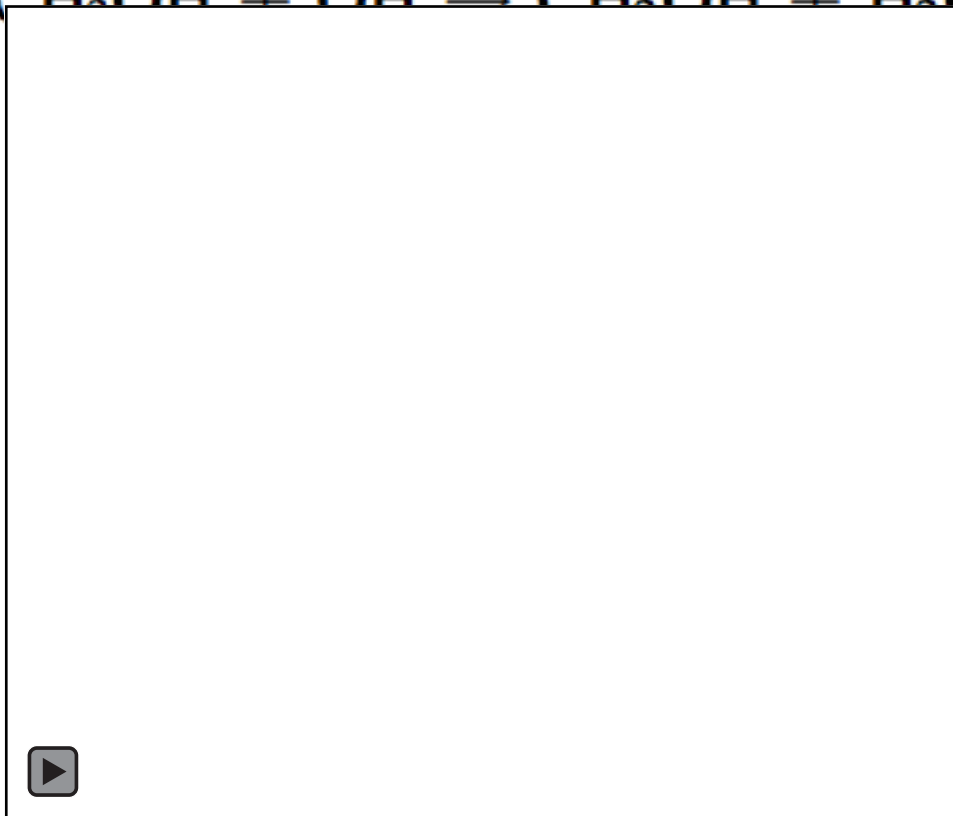
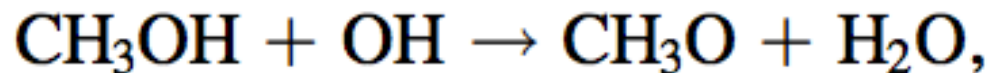


Electronic excitation energies and oscillator strengths of the quasi-linear ground triplet state $^3A'$ HCCN and singlet cyclic $^1A'$ *c*-HCCN isomers were computed using EOM-CCSDT B3LYP and ω B97-X, CIS CIS(D) //cc-pVXZ, aug-cc-pVXZ and d-aug-cc-pVXZ (X=T or Q) basis sets. Electronic excitation energies: both isomers show intense ultraviolet-visible (UV-Vis) spectra for electronic transitions with large oscillator strengths at the B3LYP, ω B97-X, and equations-of-motion coupled cluster levels.

The triplet ground state is a floppy molecule, letting a conjugation between π -electron of the CN bond to the electrons of the HC-part. Due to this, it is expected that a mixture of $^3A''$ and $^3\Sigma^-$ electronic states would produce lines in the visible region of the electromagnetic spectrum.

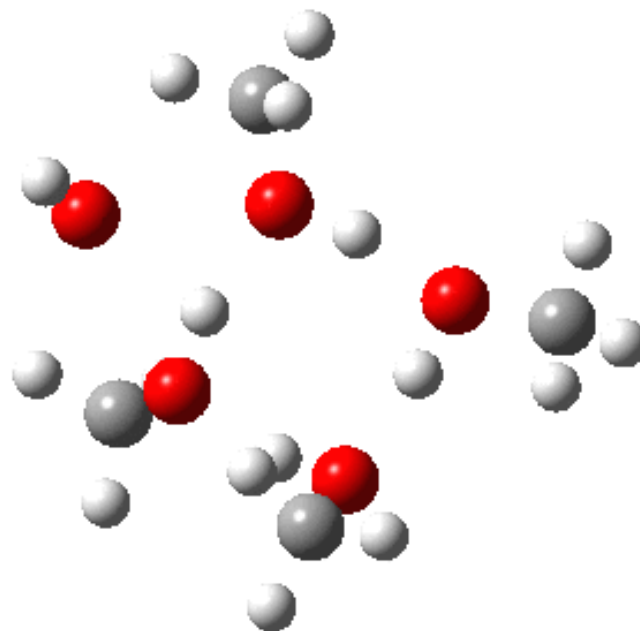
Natalia Inostroza, P. Bera, Xinchuan Huang, and Timothy J. Lee submitted June 2016

In dust we trust ??



33LYP-6-31G

BOMD
30K



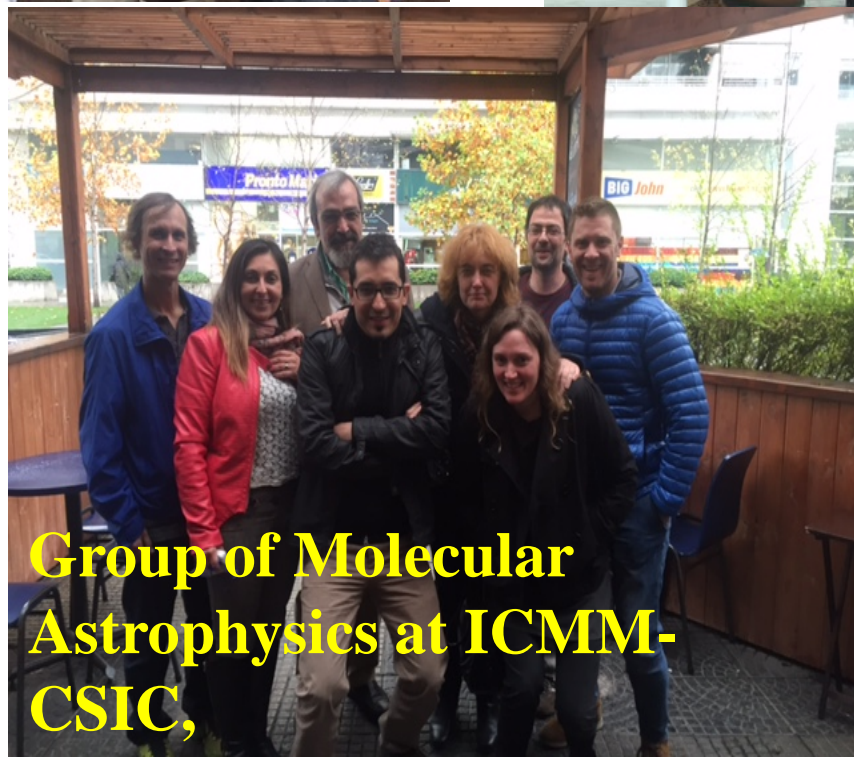
The gas-phase reaction between OH and CH₃OH is an important contributor to the formation of interstellar CH₃O. The role of grain-surface processes in the formation of CH₃O, although it cannot be fully neglected, remains controversial

Astrochemistry school, May 2016





CALAN OBSERVATORY



Group of Molecular Astrophysics at ICM- CSIC,

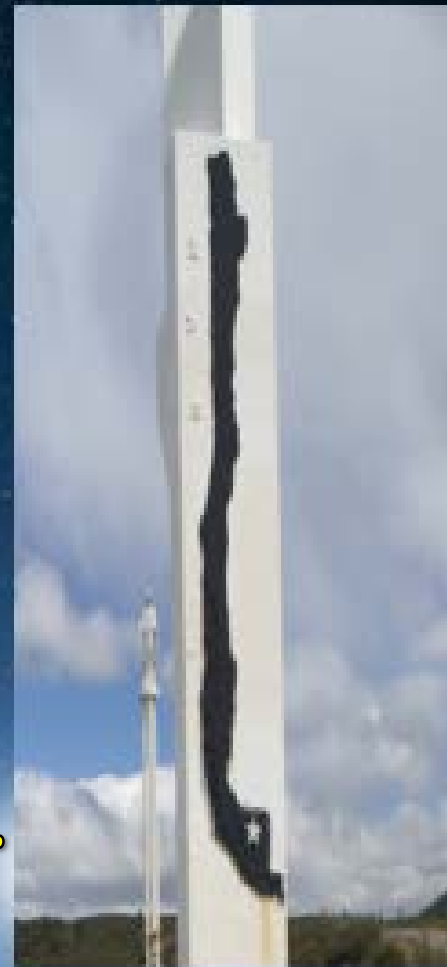
IAUS 332: Astrochemistry VII

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Thanks for you
attention



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